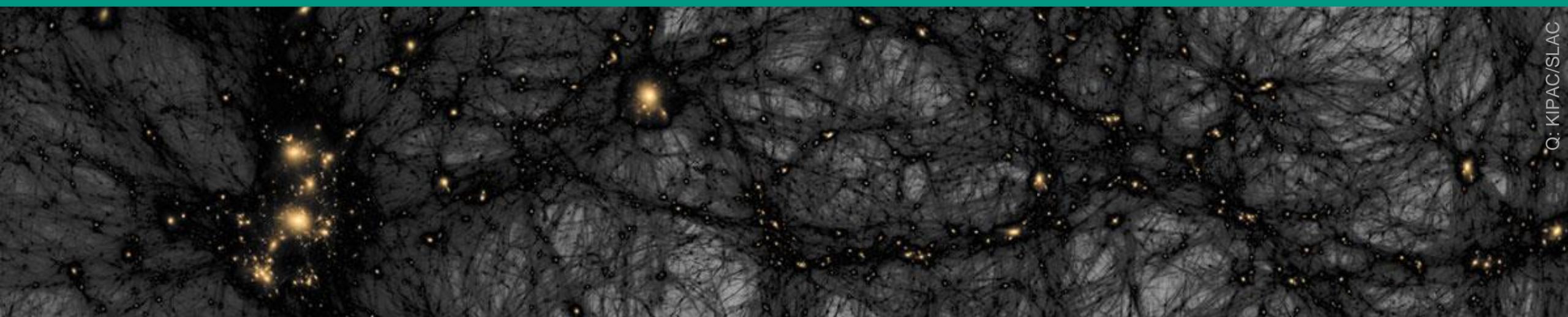


# Astroparticle physics I – Dark Matter

WS22/23 Lecture 1

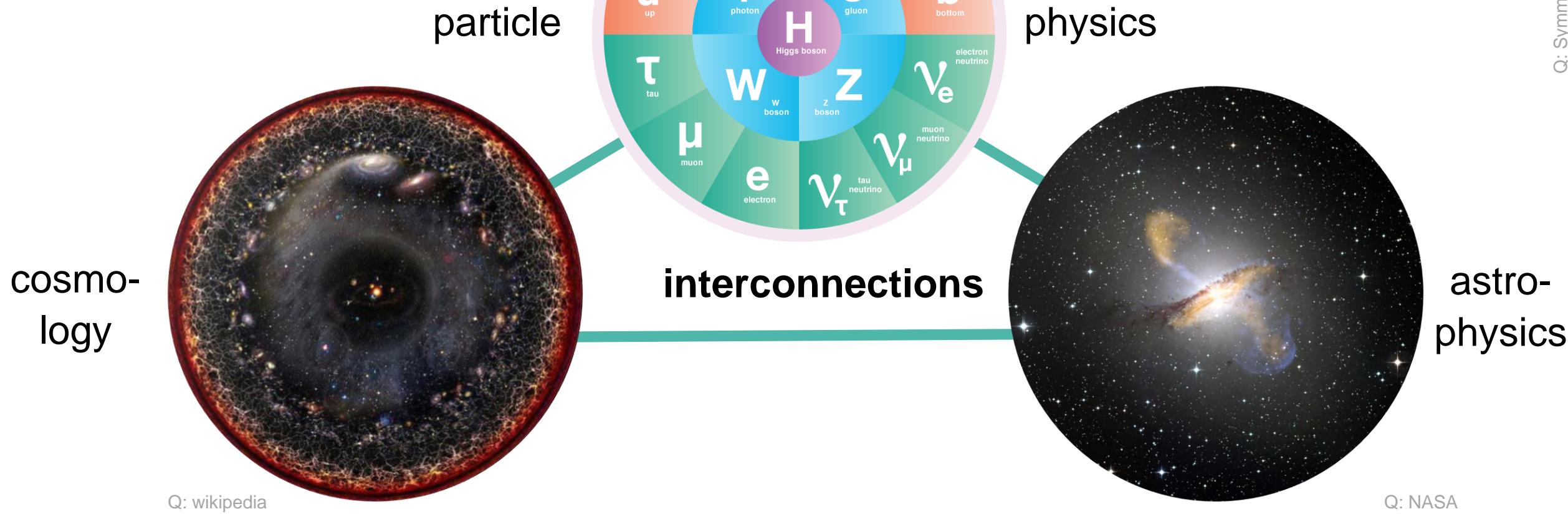
Oct. 26, 2022



Q: KIPAC/SLAC

# Astro Particle Physics – a first overview

## ■ What are our topics?



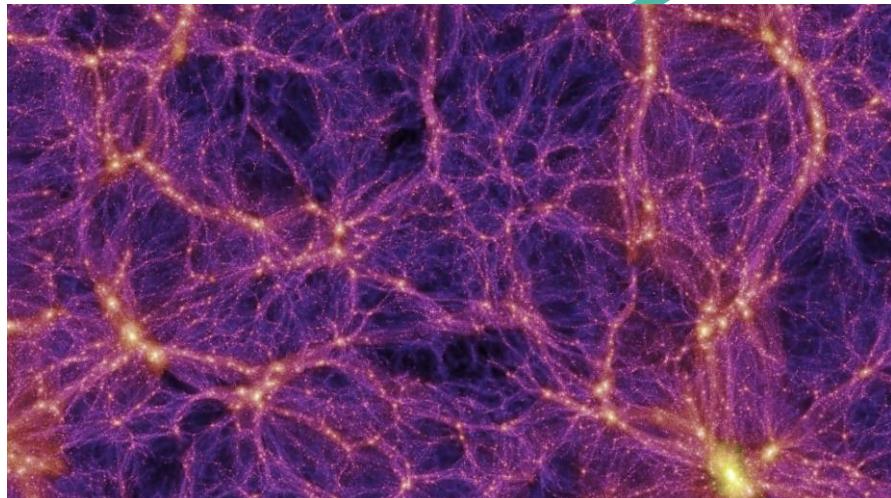
# Astro Particle Physics – a first overview

## ■ example: Dark Matter (DM)



DM candidates  
in particle physics

DM in  
cosmo-  
logy



Q: MPG

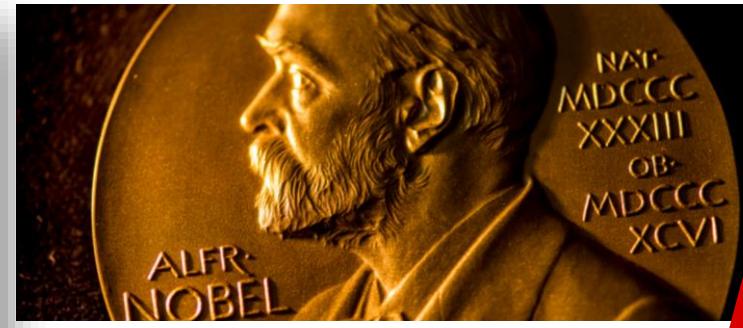


Q: wikipedia

DM in  
astro-  
physics

# Astro Particle Physics – the explosive birth event

- neutrinos: information on processes during a core-collapse supernova



**SN1987a – February '87**

Q: NASA/HST, nobelprize, TIME, SuW

# Astro Particle Physics – the central motivation

## ■ APP: deciphering the Code of the Universe



# Astro Particle Physics – the central motivation

## ■ APP: deciphering the **Code of the Universe** – here at KIT

KATRIN  
Einstein-Telescope  
GridKa



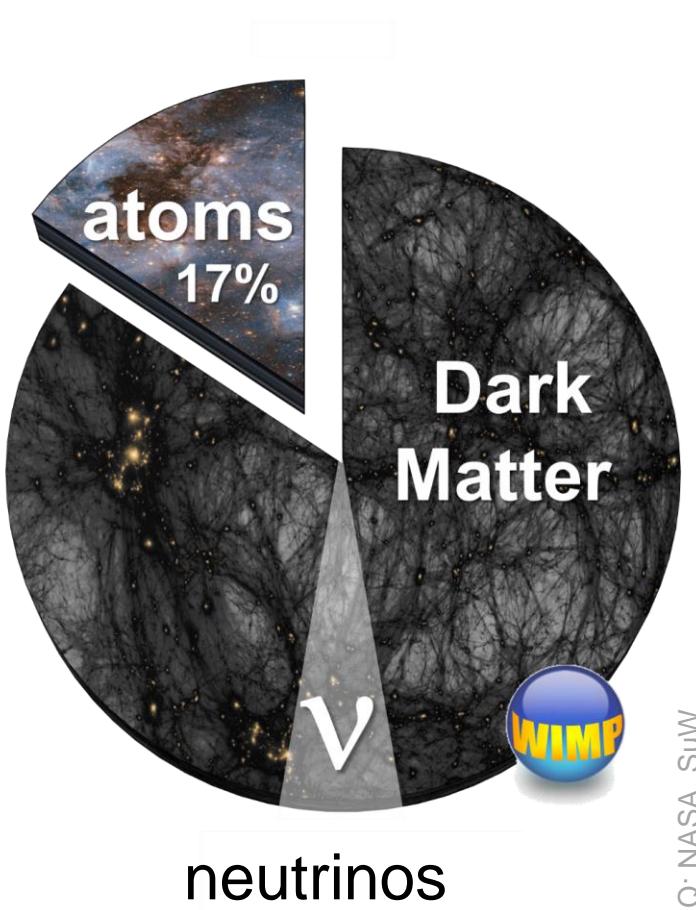
Auger  
IceCube  
XENON-nT

The collage features several images and text elements:

- Three square images at the top right show the interior of a particle accelerator, a long telescope array in space, and a globe surrounded by particle tracks.
- A large central title "CODE DES UNIVERSUMS" is written in white, with the letter "U" having a yellow outline.
- To the left of the title, the text "our experiments" is displayed.
- Below the title, there are three more images: one showing a particle interaction with Earth, one showing a detector in a snowy landscape with a signal waveform, and one showing a detector in space with a figure standing next to it.

# Astro Particle Physics – the central motivation

## ■ neutrinos & neutralinos: from the Big Bang & the universe



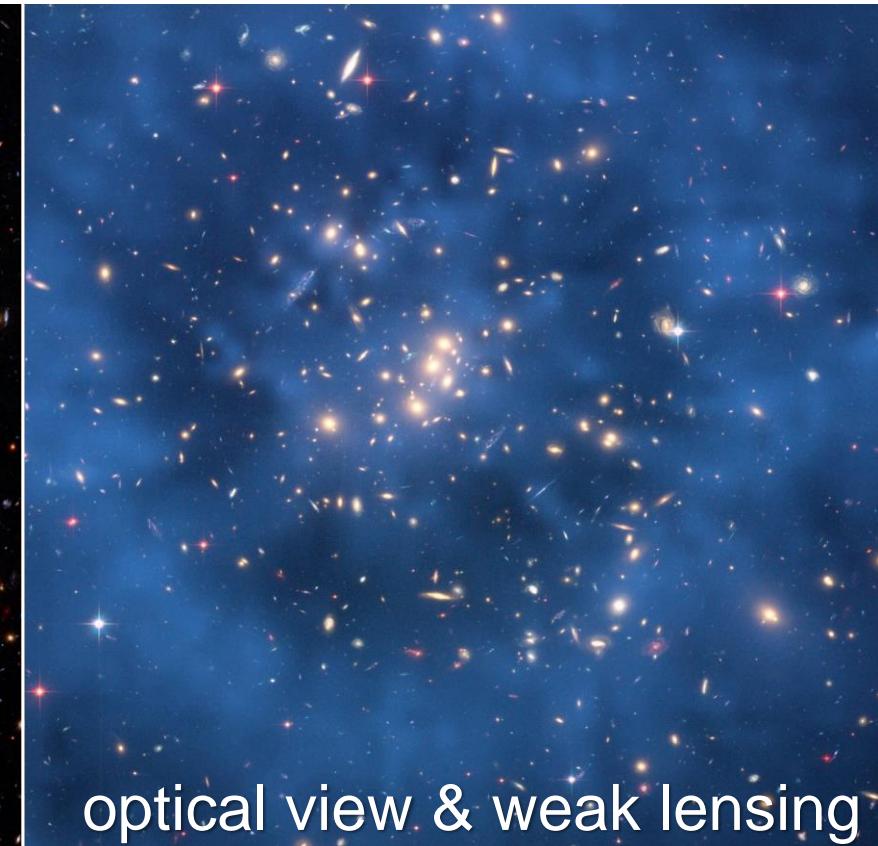
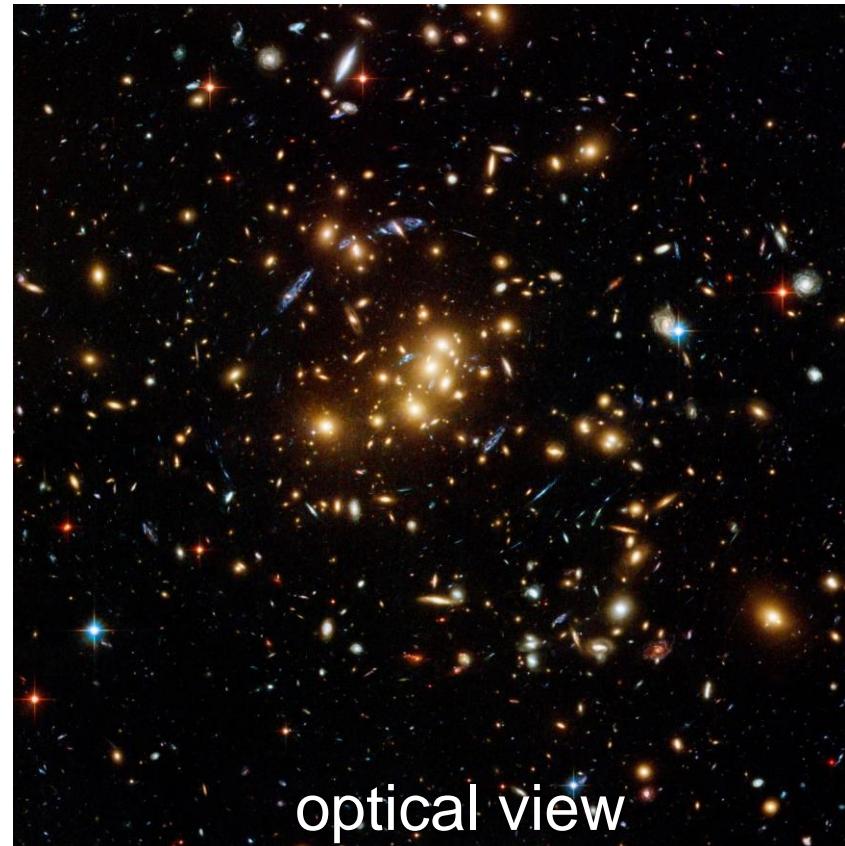
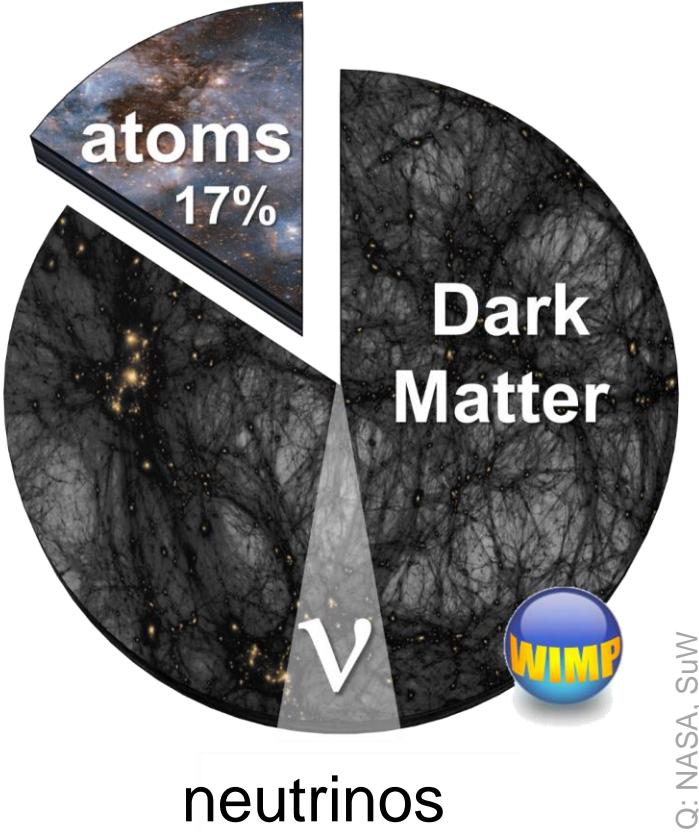
- massive **neutrinos** as **Hot Dark Matter - HDM**
- massive **neutralinos** as **Cold Dark Matter - CDM**



# Astro Particle Physics & Cosmology: perfect match

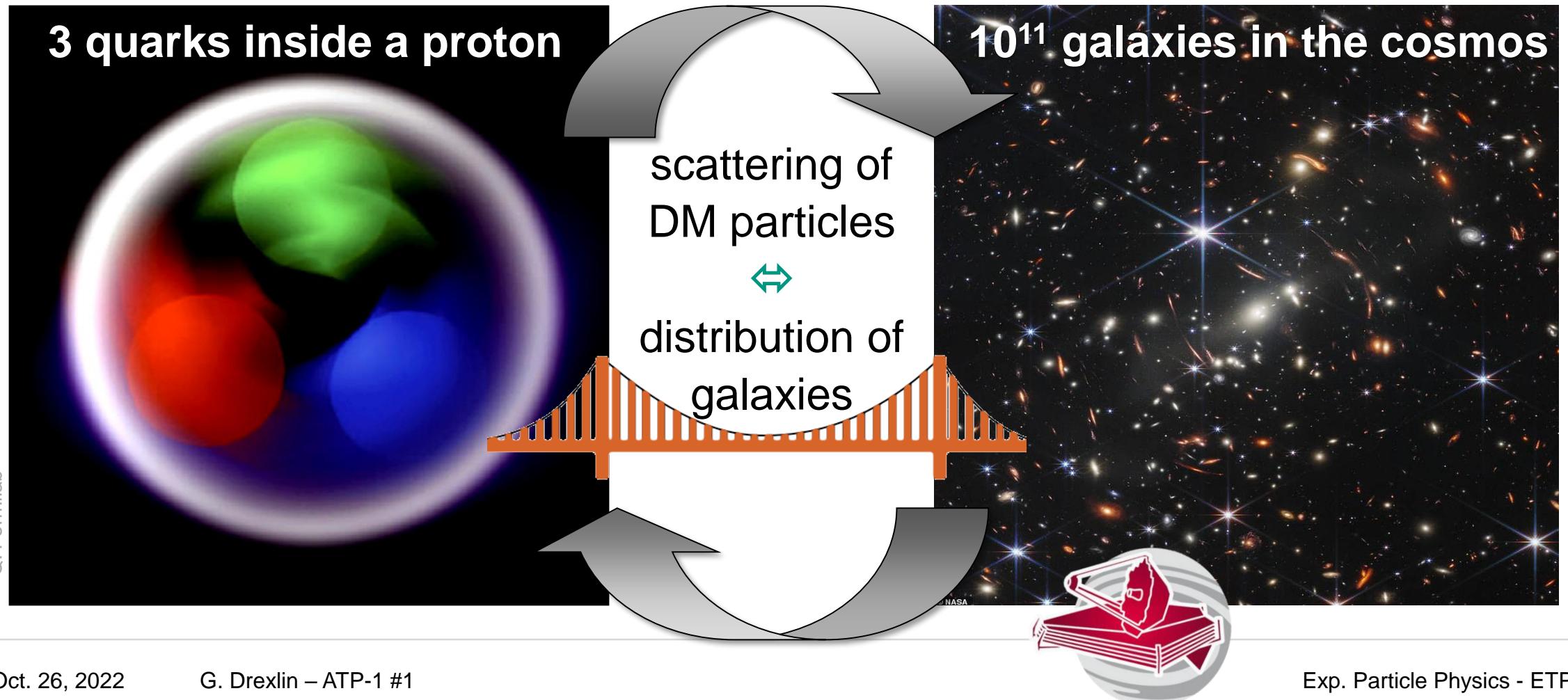
- Dark Matter: maps by gravitational lenses, see 'Introduction to Cosmology'

Cosmology lectures: Tuesdays 11:30-13:00. kl. HS. A



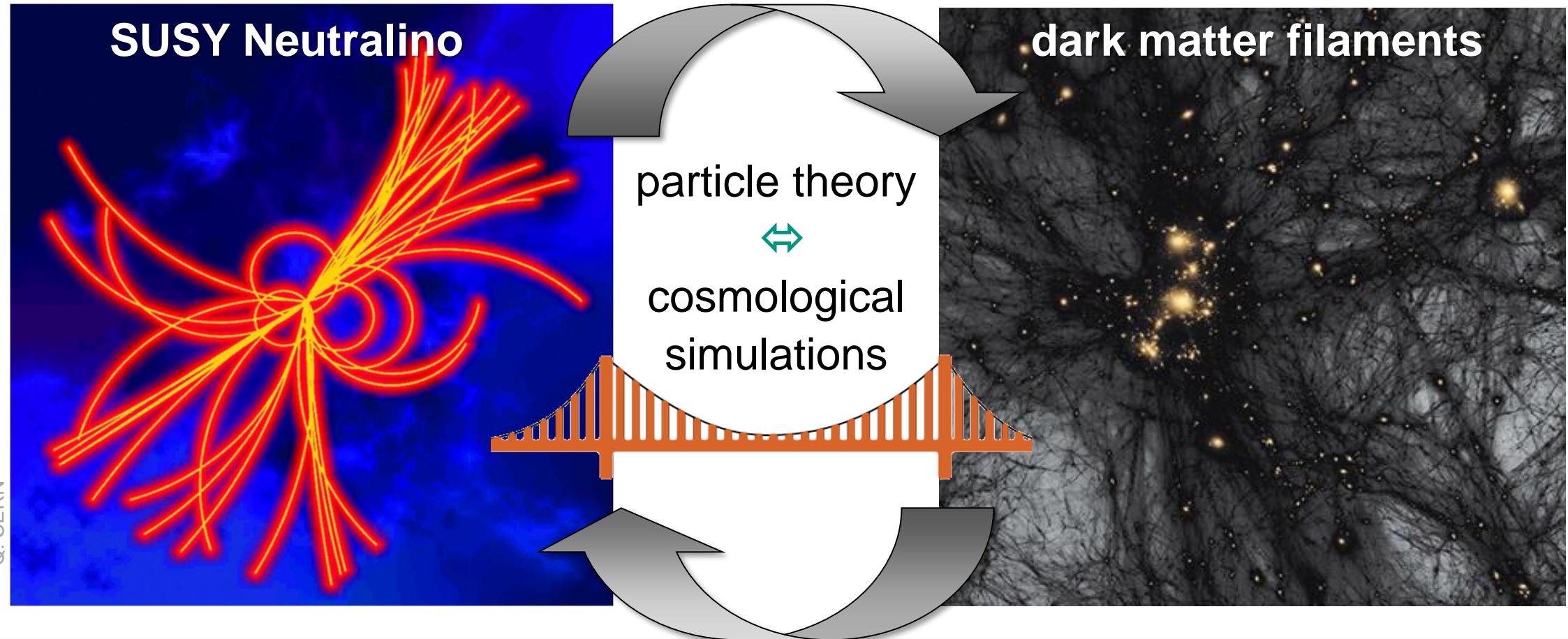
# Astro Particle Physics – cross connections #1

■ ...from the **quark level** to the entire **cosmos**...



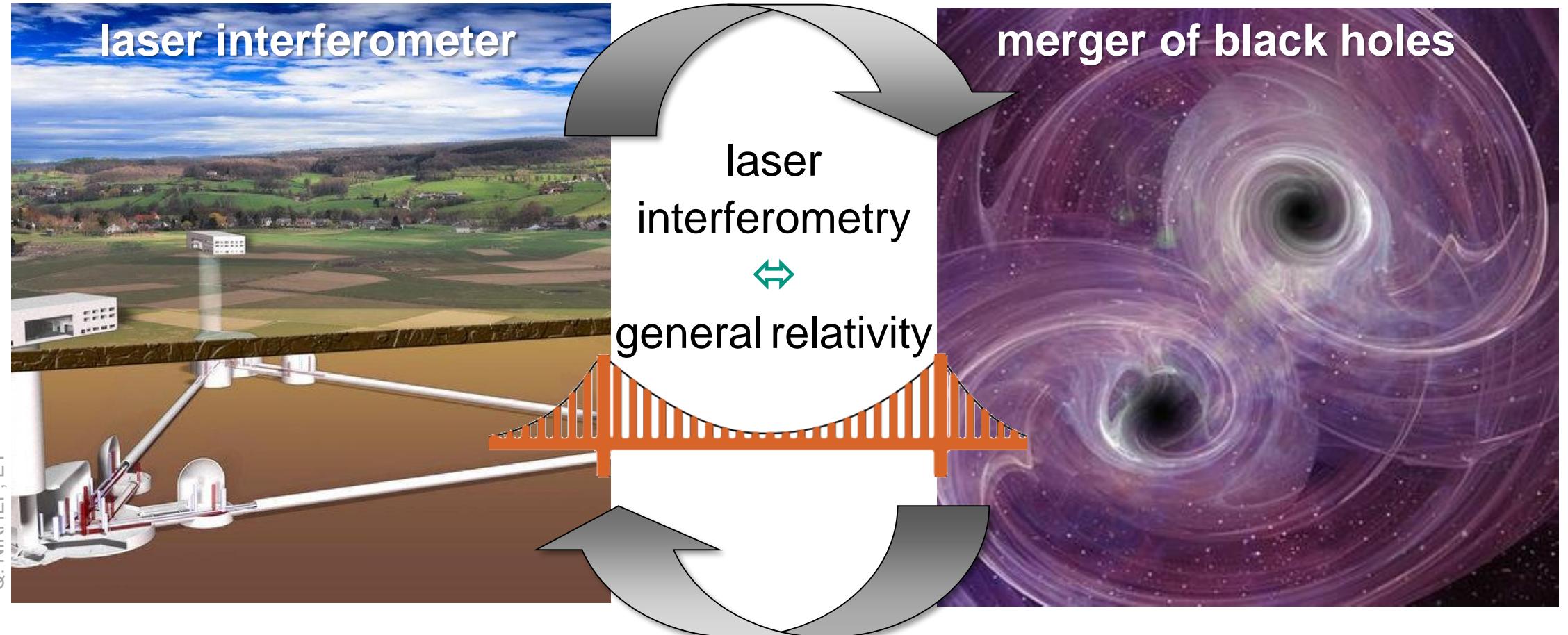
# Astro Particle Physics – cross connections #2

■ ...from **neutralinos** to **dark matter** in the **cosmos**...



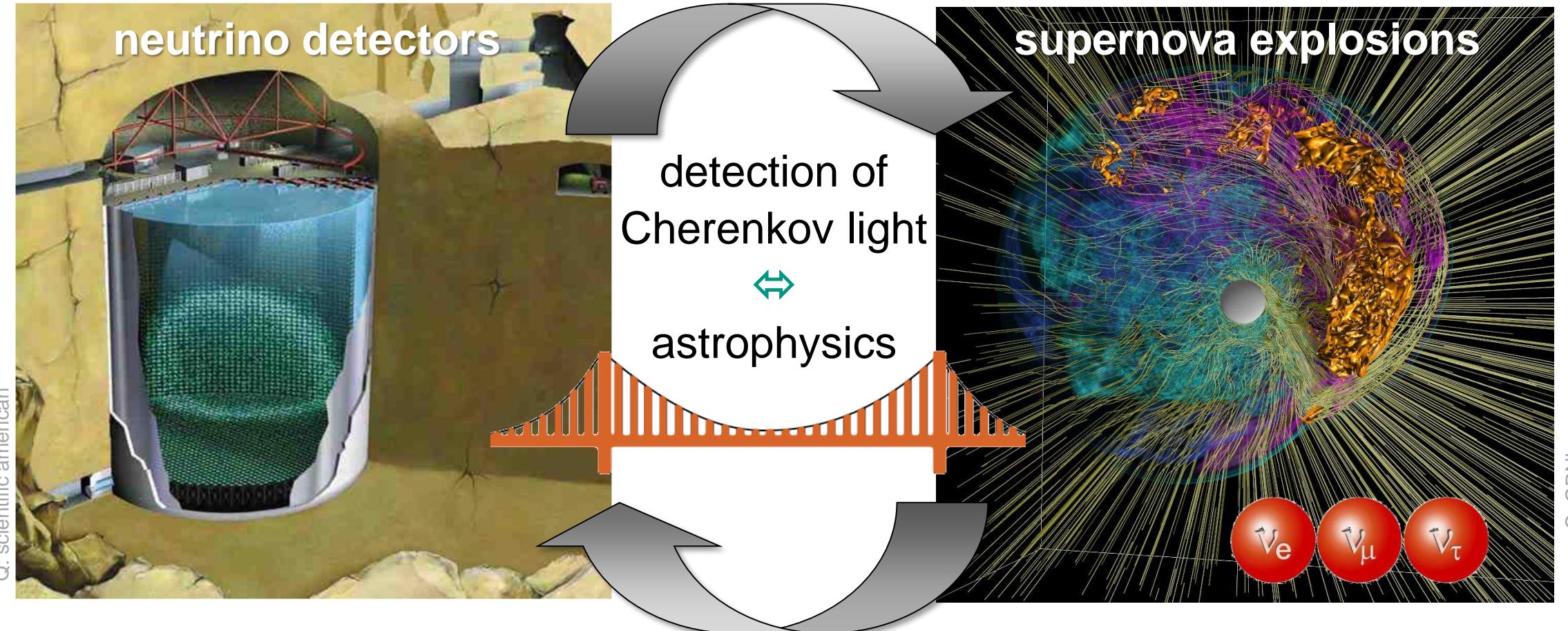
# Astro Particle Physics – cross connections #3

- ...from gravitational waves to black holes & strong gravity (GR\*)



# Astro Particle Physics – cross connections #4

- ...from neutrino detectors to core-collapse supernovae

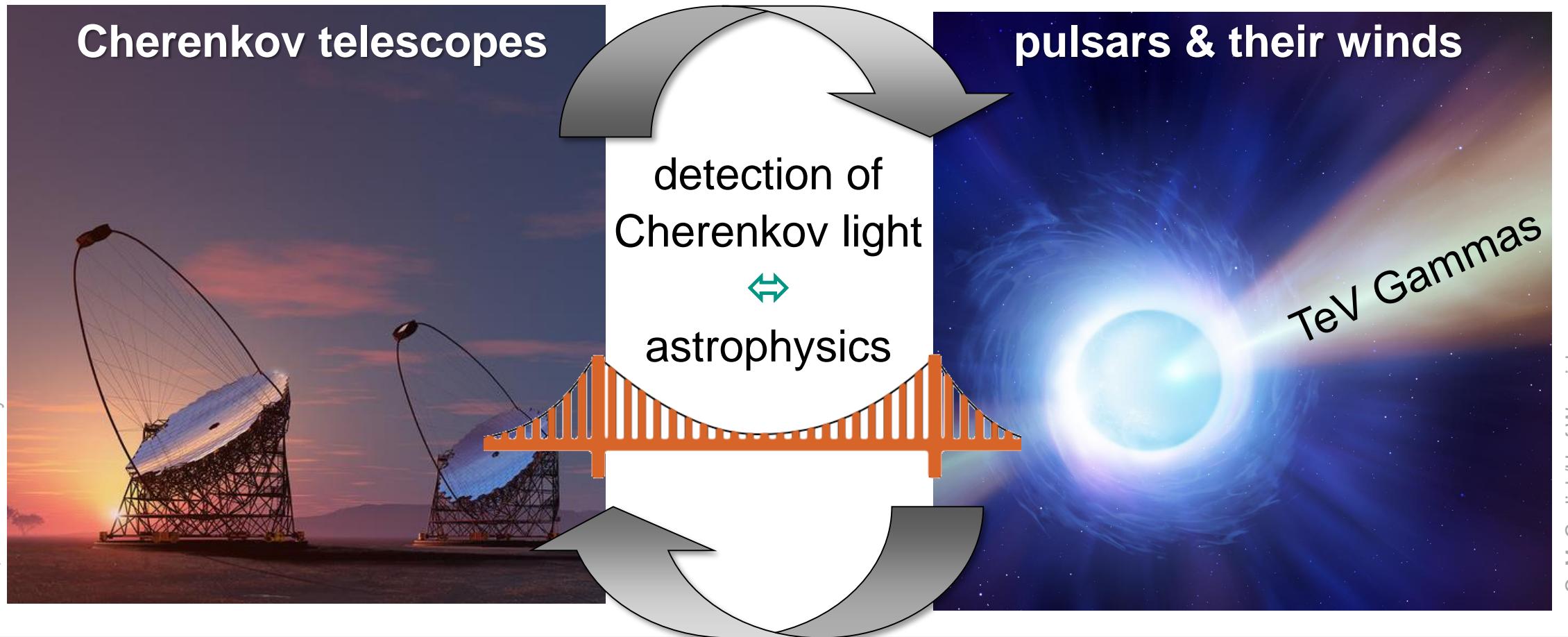


Q: scientific american

Q: ORNL

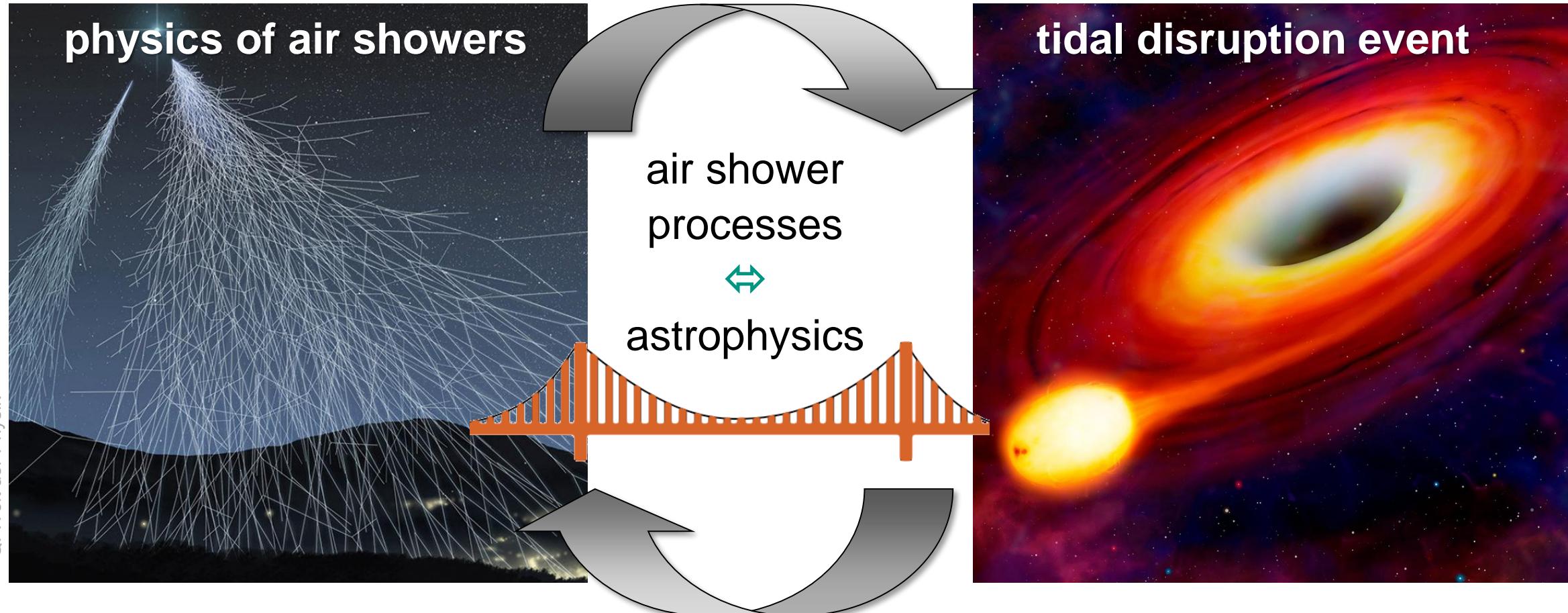
# Astro Particle Physics – cross connections #5

- ...from TeV-Cherenkov telescopes to pulsar wind nebulae



# Astro Particle Physics – cross connections #6

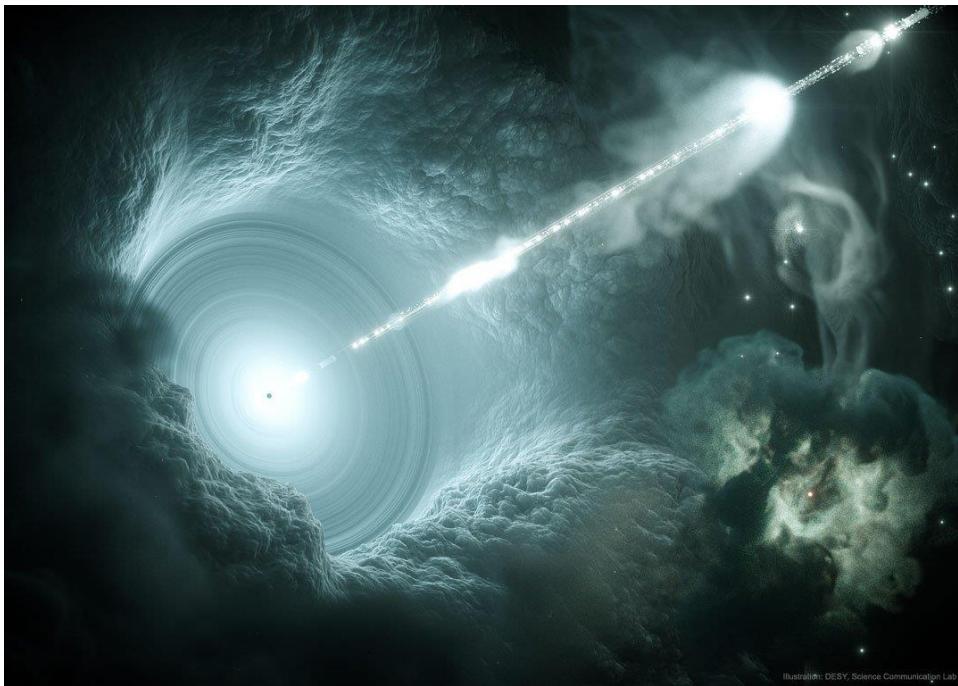
- ...from **air shower arrays** to **cosmic super accelerators**



# Astro Particle Physics – recent breakthroughs

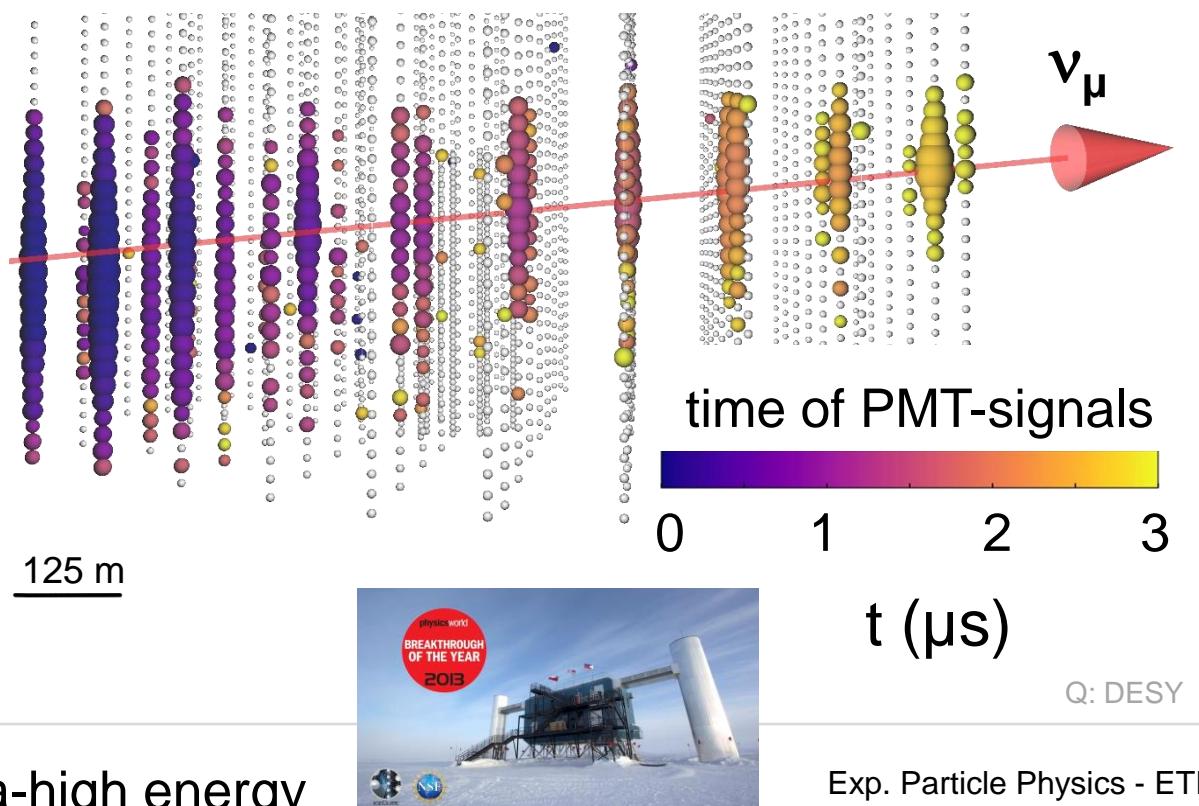
## ■ ...from ultra-high energy neutrinos to blazars...

‘Texas source’: TXS 0506+056  
an active supermassive black hole



Blazars as sources for UHE\* ν's

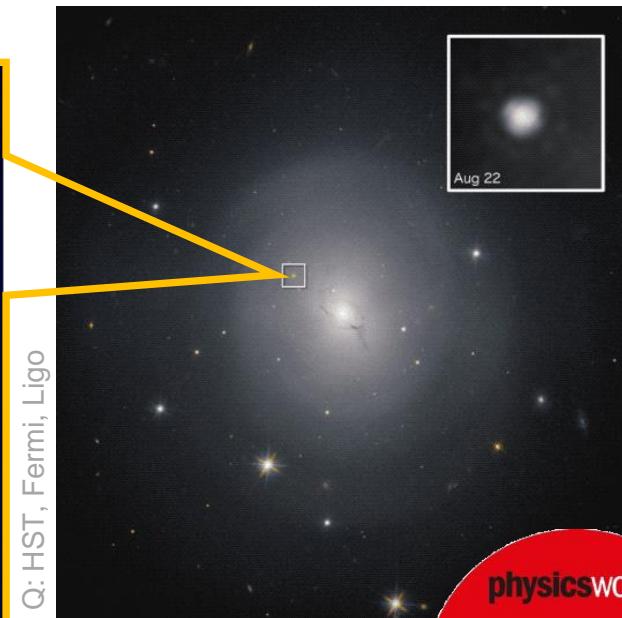
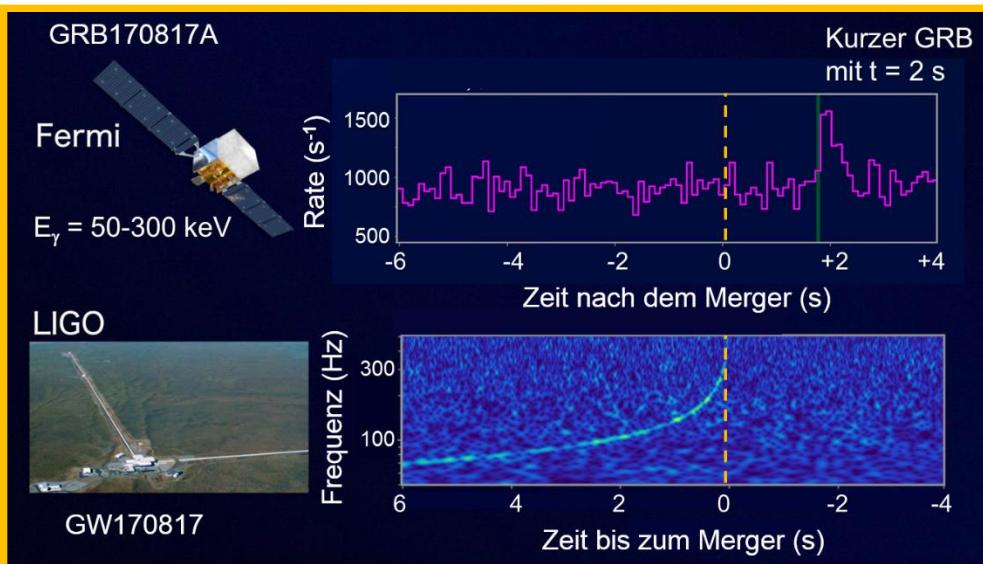
Sept. 22, 2017: IceCube observes  
an event with  $E = 290 \text{ TeV}$



# Astro Particle Physics – recent breakthroughs

## ■ ...from gravitational waves to GRB\*s...

merger process of two neutron stars observed *live*



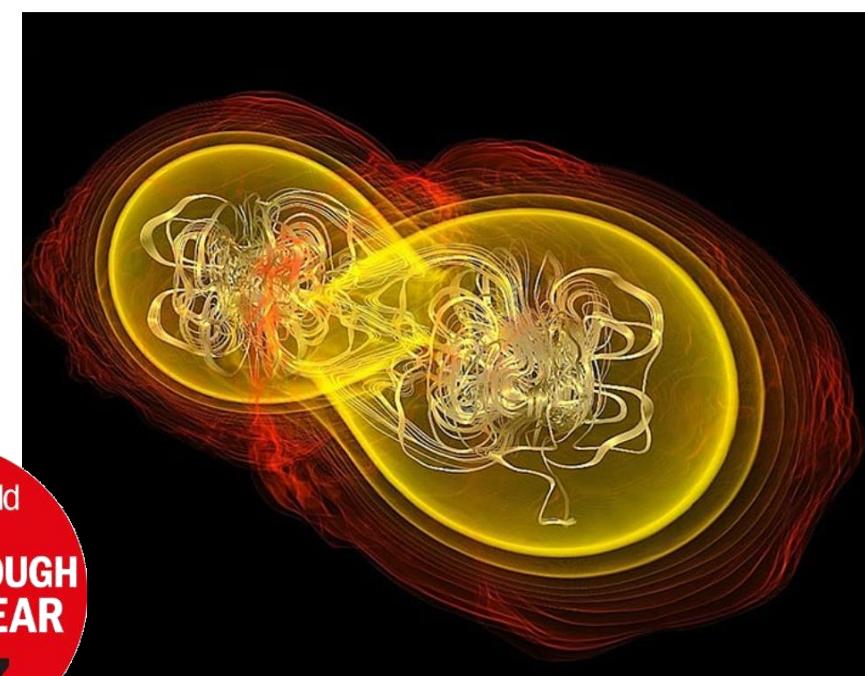
Q: HST, Fermi, Ligo



news on the origin of  
*heavy elements*

Shrouded in mystery: artist's impression of merging neutron stars

Q: physics world



# Astro Particle Physics – a very recent observation

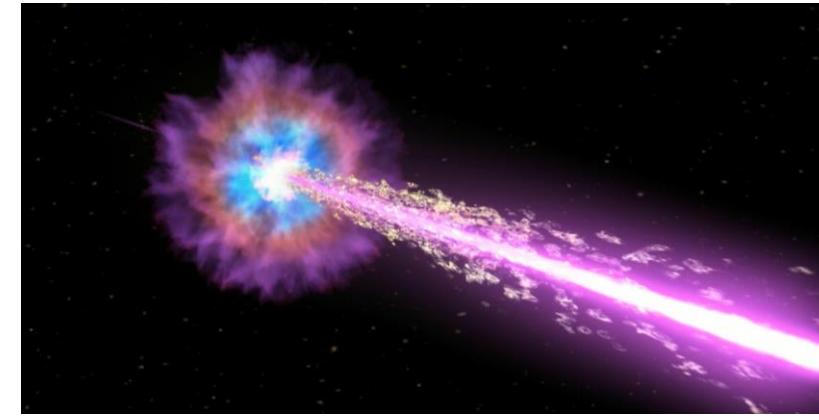
- ...a unique record GRB occurred only ~ two weeks (October 9, 2022) ago...

- collapse of a massive star to a black hole

generation of a very long burst (~10 h) of gamma rays with peak energies up to 14 TeV

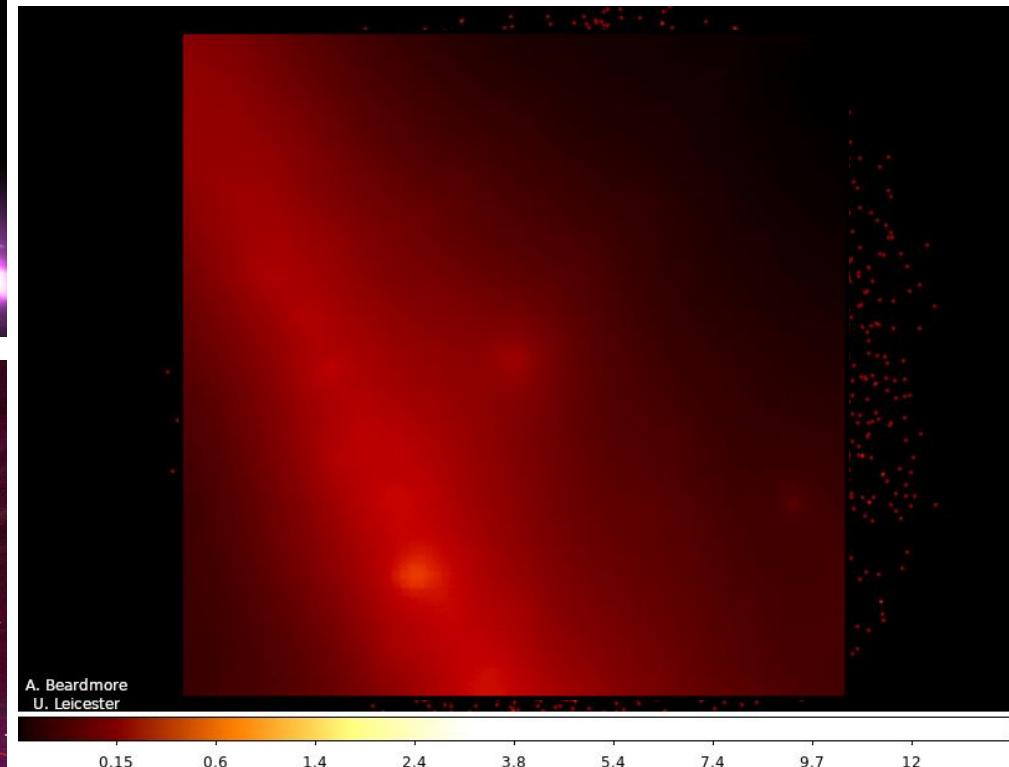
**HOT TOPIC**

GRB occurred at a distance  $d = 2$  Gpc  
here: light reflection at dust grains in between



OCTOBER 15, 2022

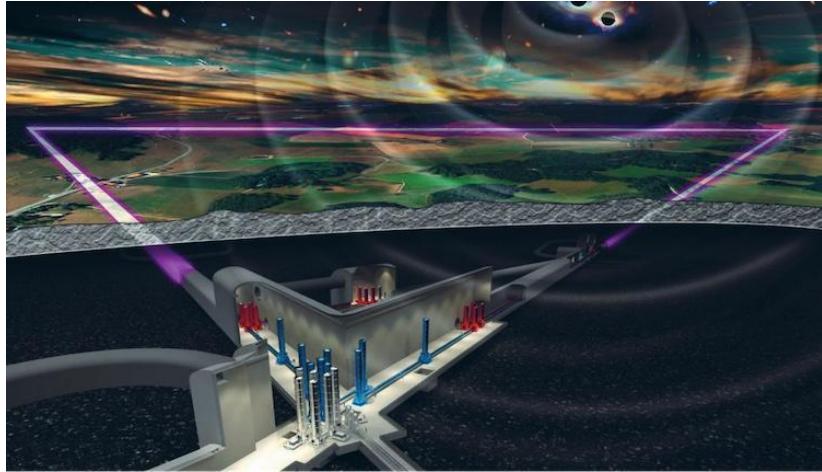
Record-breaking gamma-ray burst possibly most powerful explosion ever recorded



# Astro Particle Physics – future breakthroughs

## ■ DZA – Deutsches Zentrum für Astrophysik: a new centre in the Lausitz

- bundling the activities of German astroparticle physics

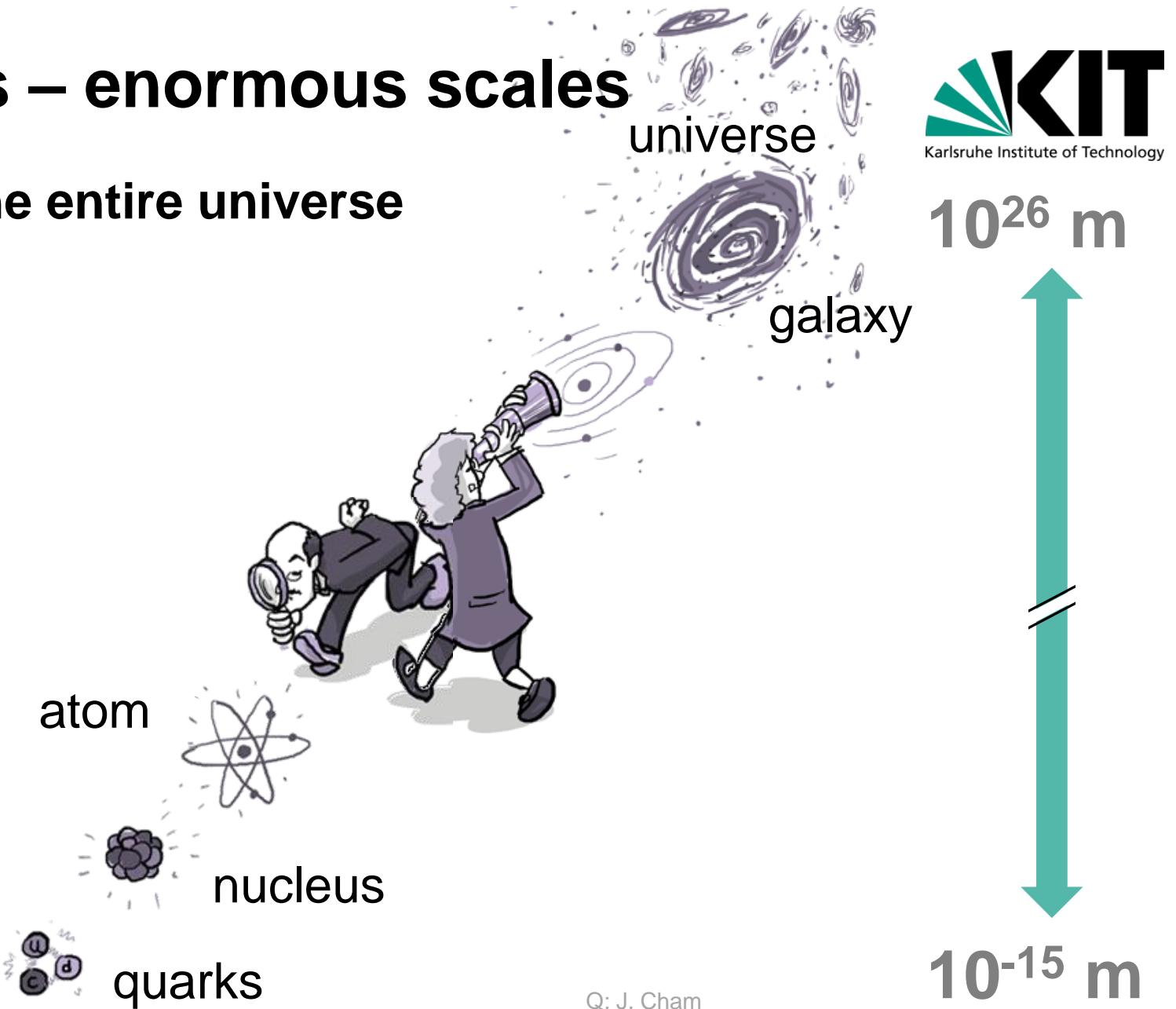
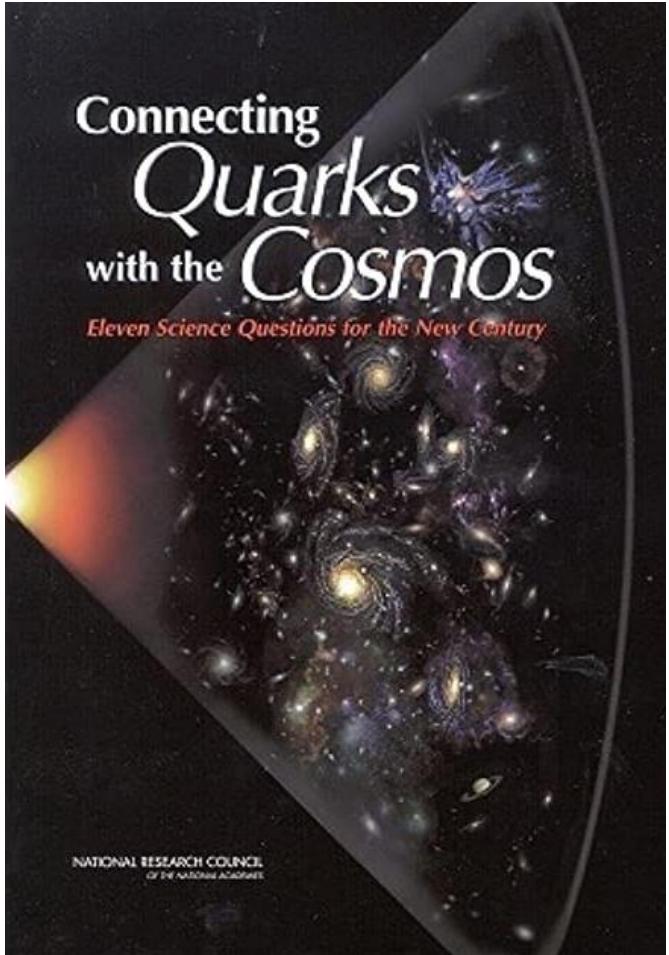


The screenshot shows the DZA website's homepage. At the top right are language links "DE | EN". The main header features the "DZA" logo inside a white triangle. To the right is a large circular image of a galaxy or nebula. Below the header is a navigation bar with links: Mission, Center, Research and transfer, Partner, and News. A text box at the bottom contains the following message:

Wir haben es geschafft: Das Deutsche Zentrum für Astrophysik - Forschung, Technologie, Digitalisierung. (DZA) kommt in die sächsische Lausitz. Hier geht's zu der Pressemitteilung:

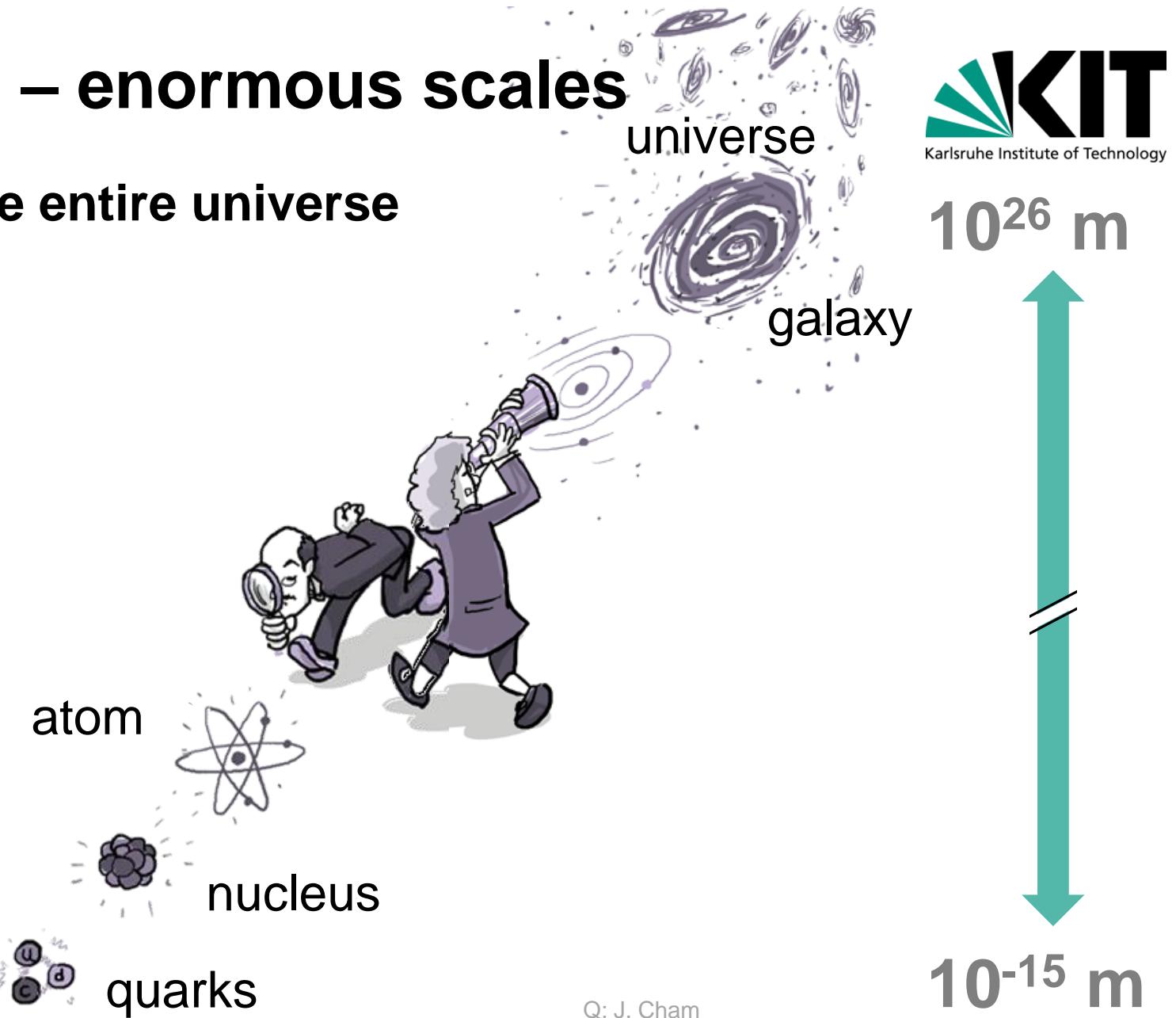
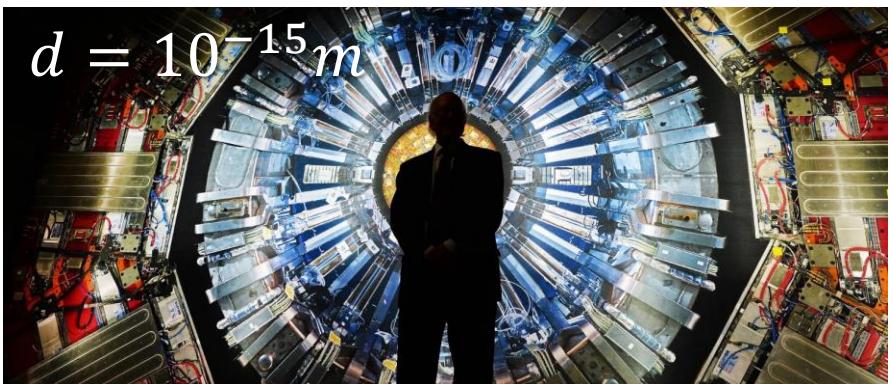
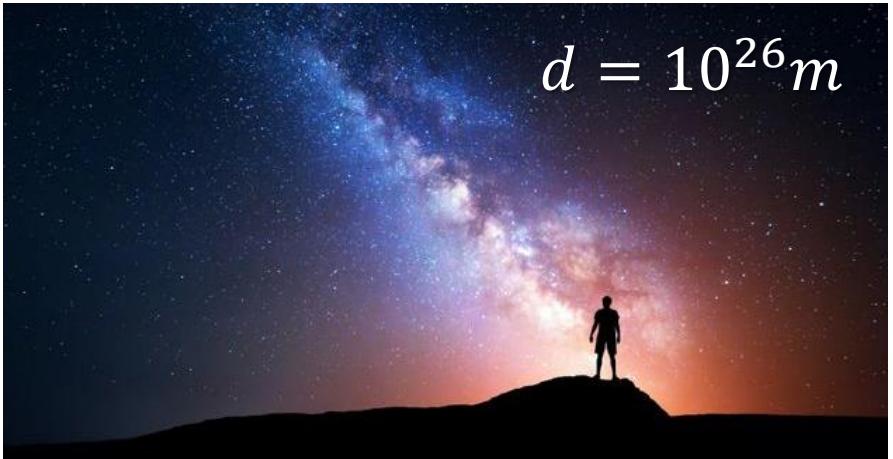
# Astro Particle Physics – enormous scales

- from the quark scale to the entire universe



# Astro Particle Physics – enormous scales

■ from the quark scale to the entire universe

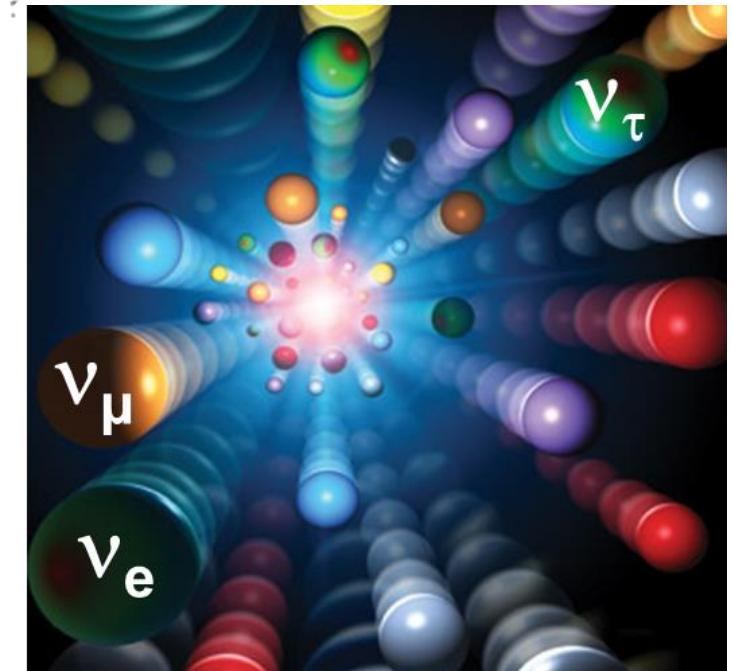
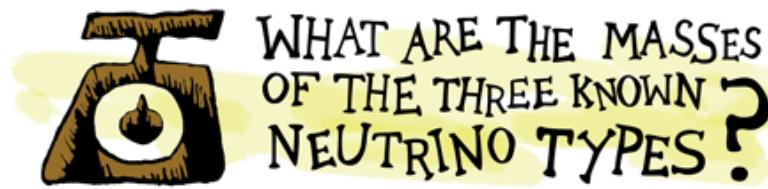


# Astro Particle Physics – open questions



## ■ what is the nature of neutrinos?

- we can find out by **observing their role** in the cosmos!



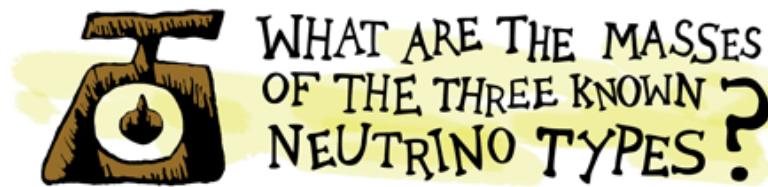
- using our knowledge of cosmic evolution  
galaxy formation

Q: J. Cham

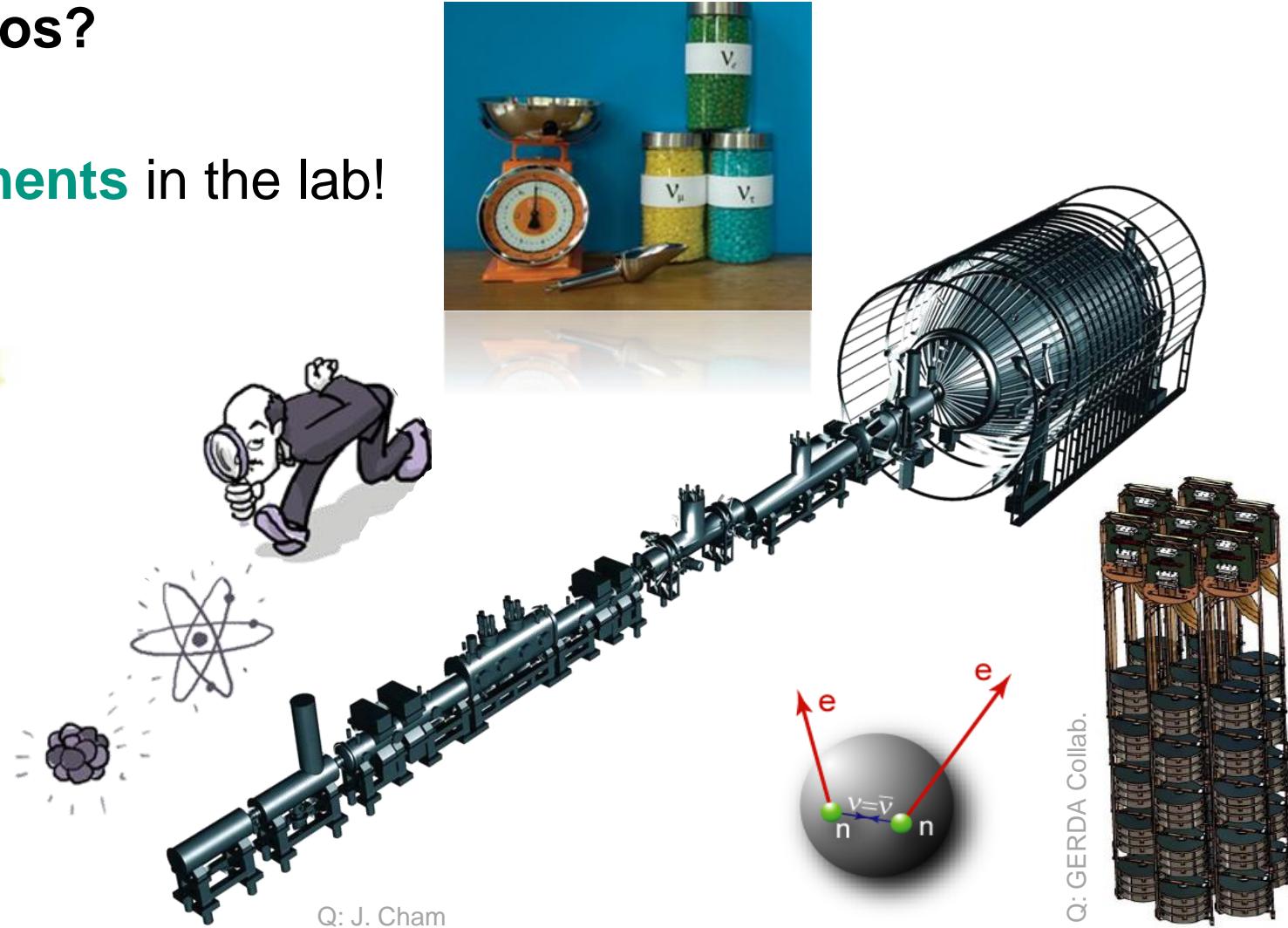
# Astro Particle Physics – open questions

## ■ what is the nature of neutrinos?

- we can find out by doing **experiments** in the lab!



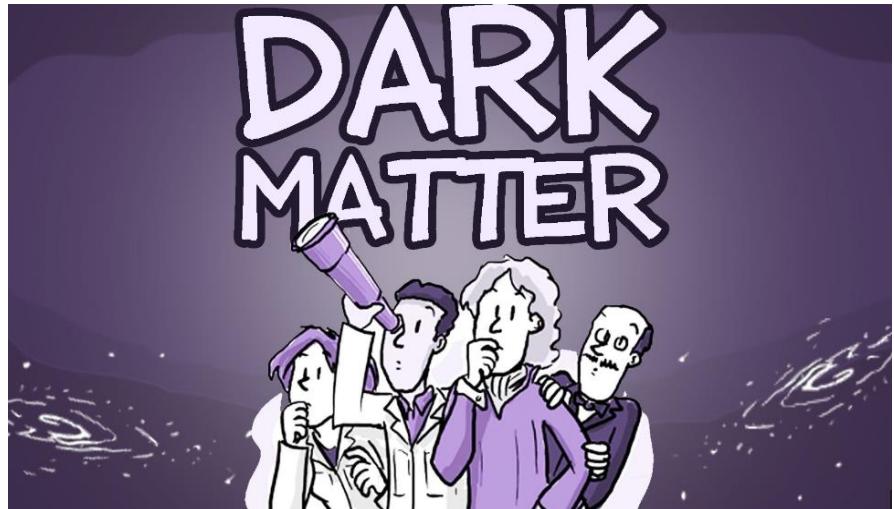
- using our knowledge of  **$\beta$ -decay theory**



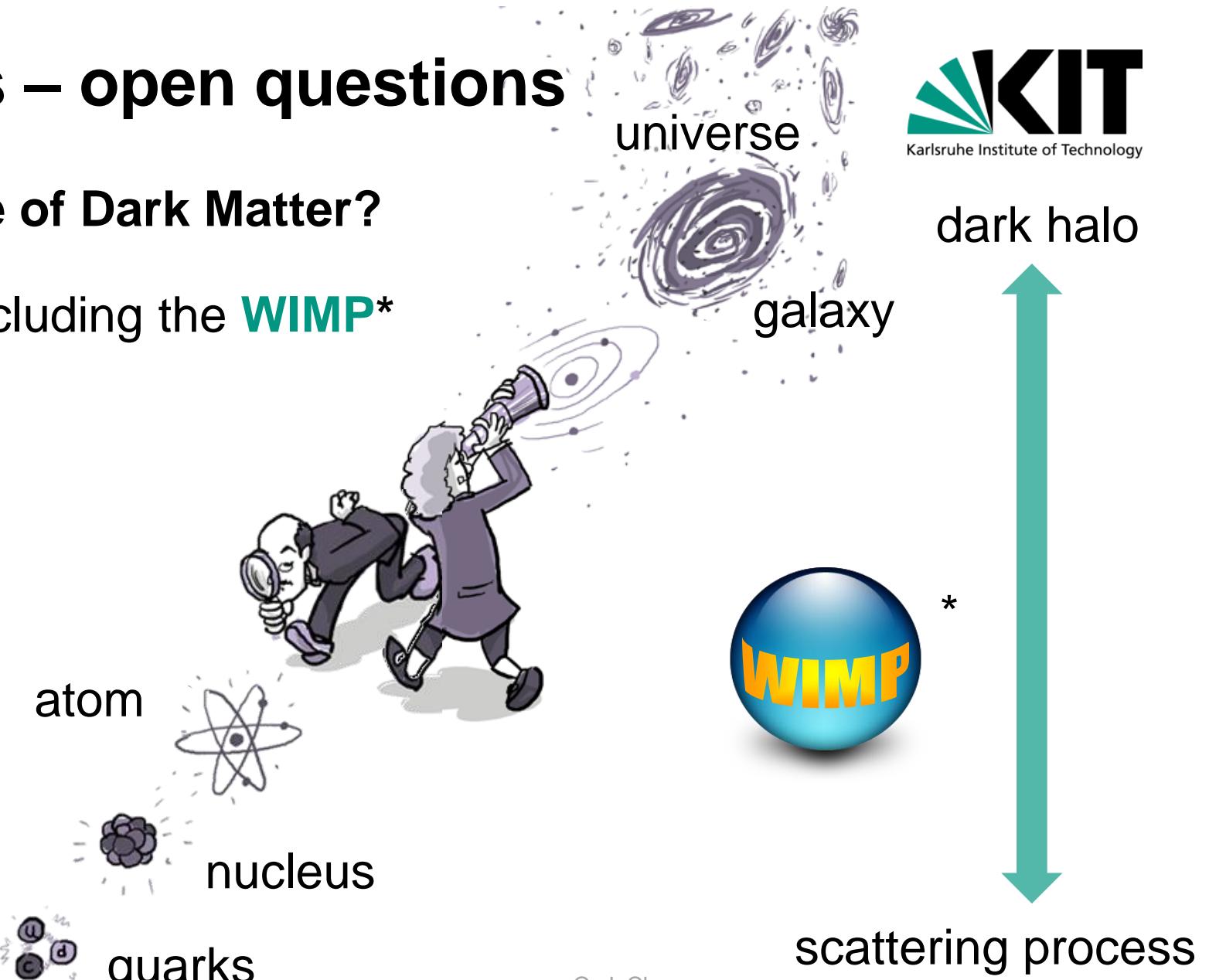
# Astro Particle Physics – open questions

## ■ what is the particle nature of Dark Matter?

- there are many candidates, including the **WIMP\***



- how can we detect them?
- how to measure their mass?

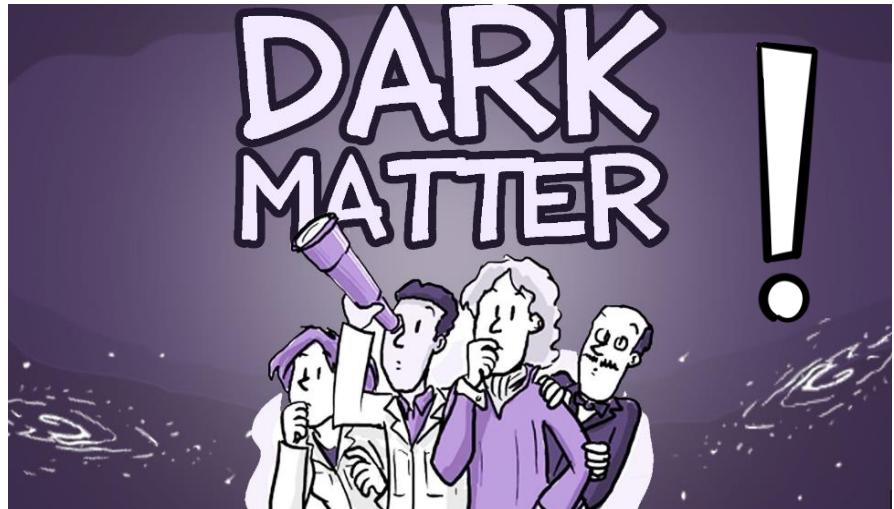


Q: J. Cham

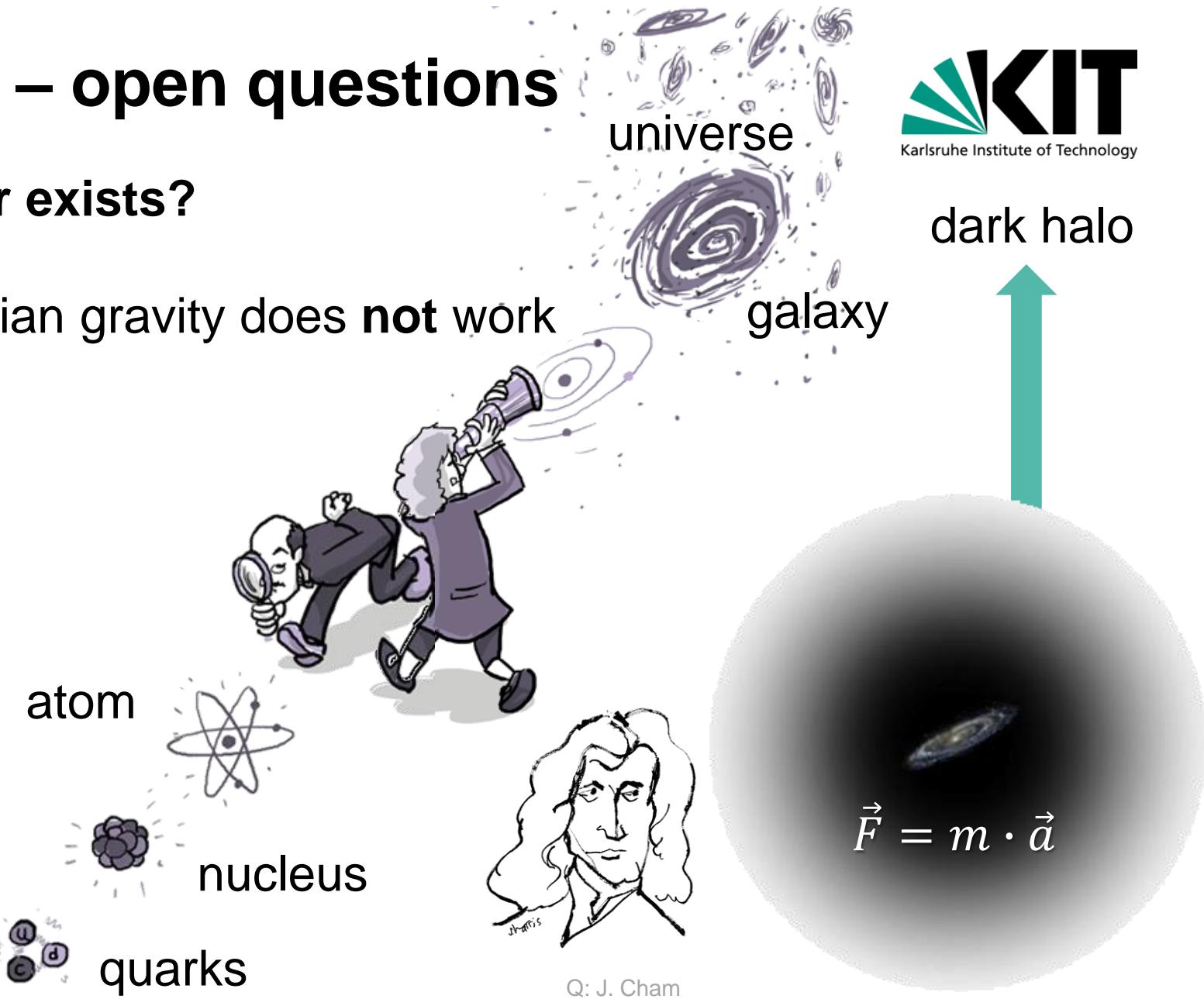
# Astro Particle Physics – open questions

## ■ Are we certain Dark Matter exists?

- yes, we are! To modify Newtonian gravity does **not** work



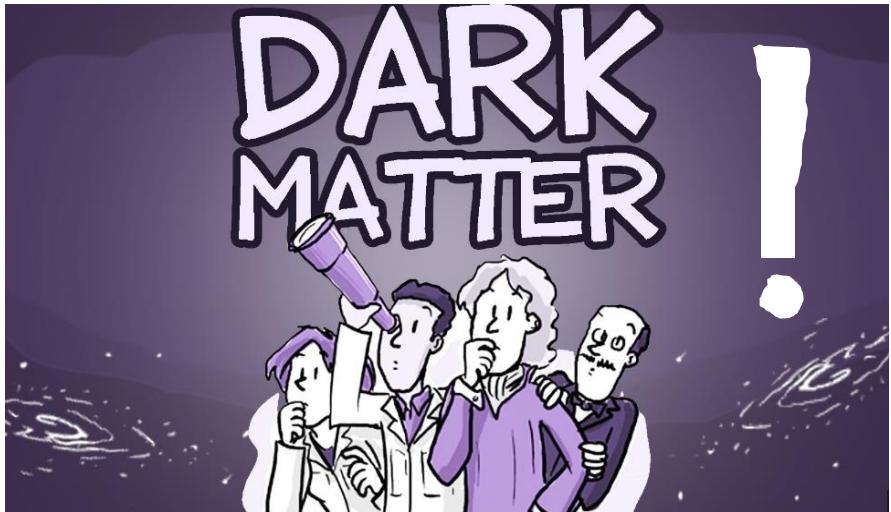
- Dark matter halos surround all galaxies, plus other evidences



# Astro Particle Physics – open questions

## ■ what is the particle nature of Dark Matter?

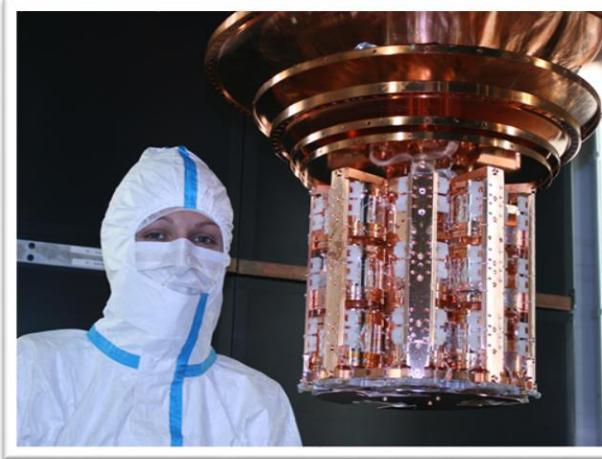
- we can find out via **detecting it here on Earth!**



- using novel advanced detectors
- background reduction methods



Q: J. Cham



Q: CRESST

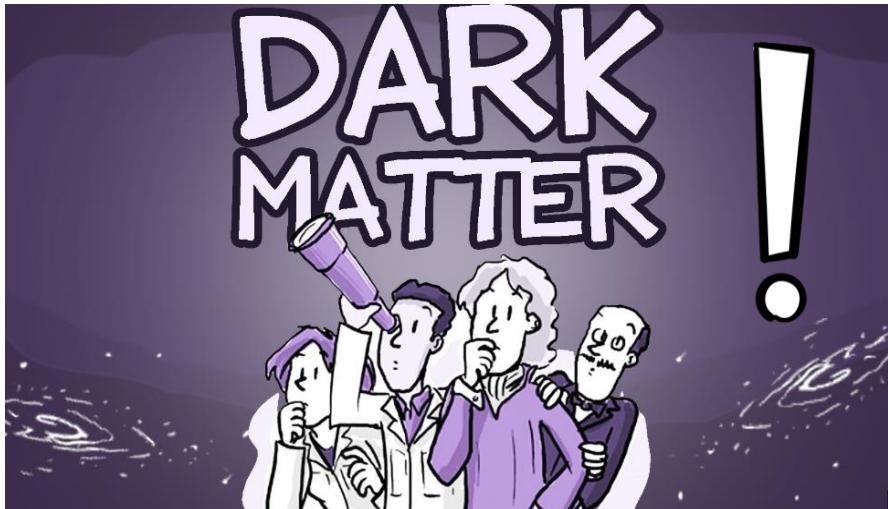


Q: XENON/DARWIN

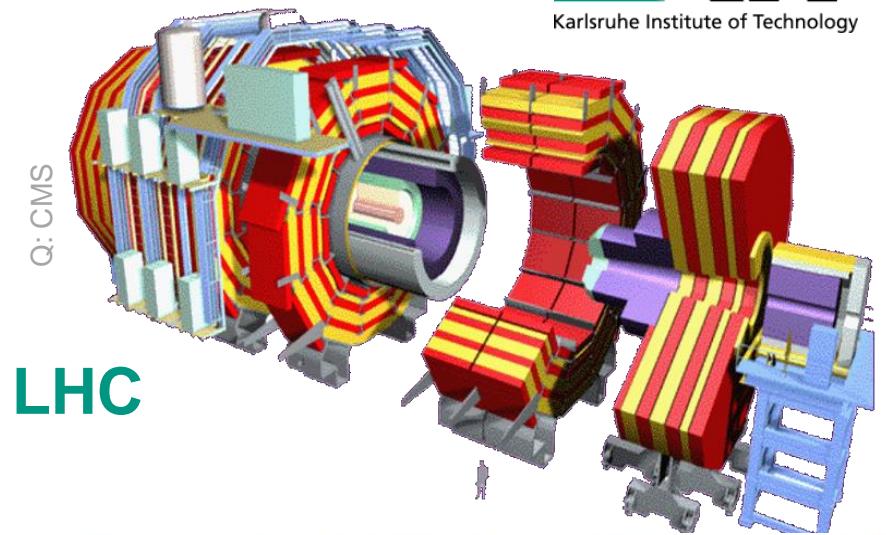
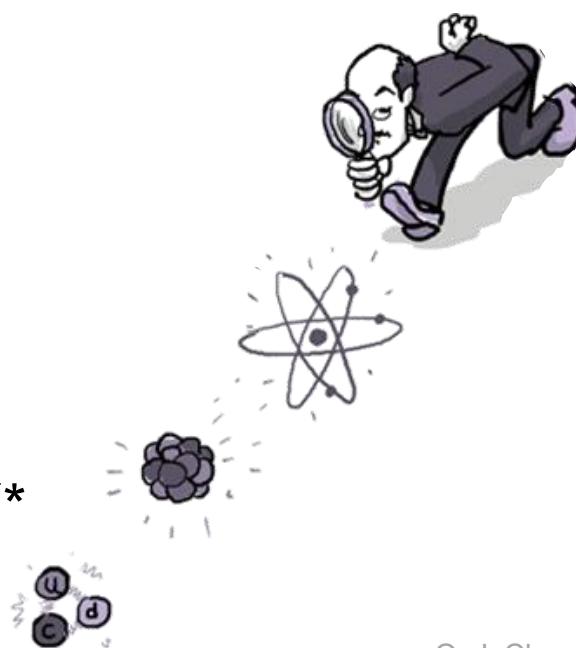
# Astro Particle Physics – open questions

## ■ what is the particle nature of Dark Matter?

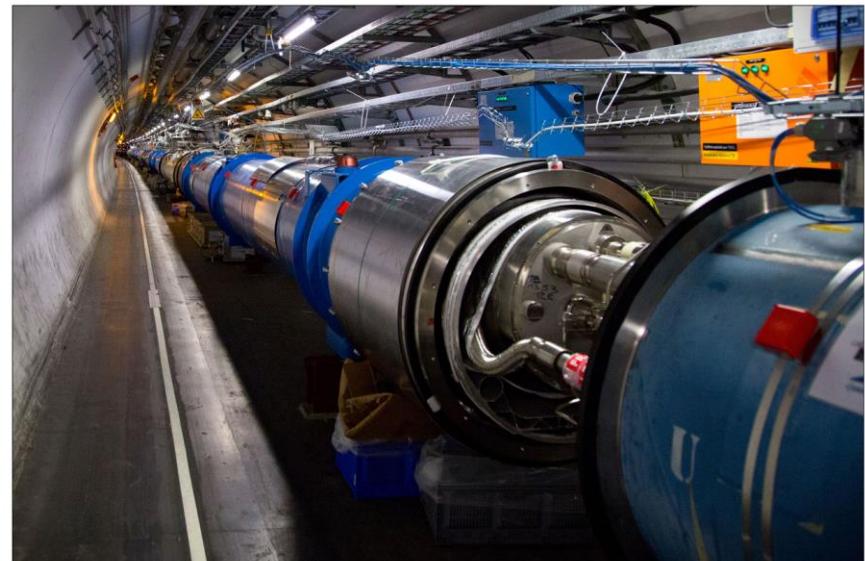
- we can find out by **producing it** via an accelerator!



- using our knowledge of SUSY\*
- using our knowledge of LHC



LHC

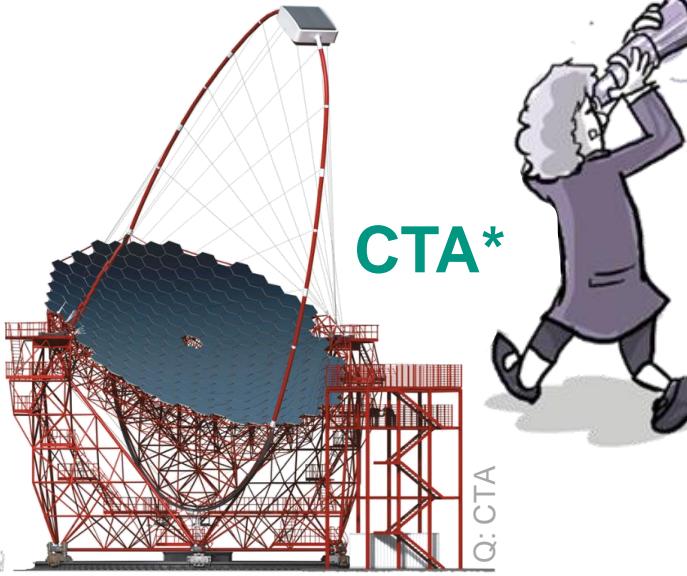
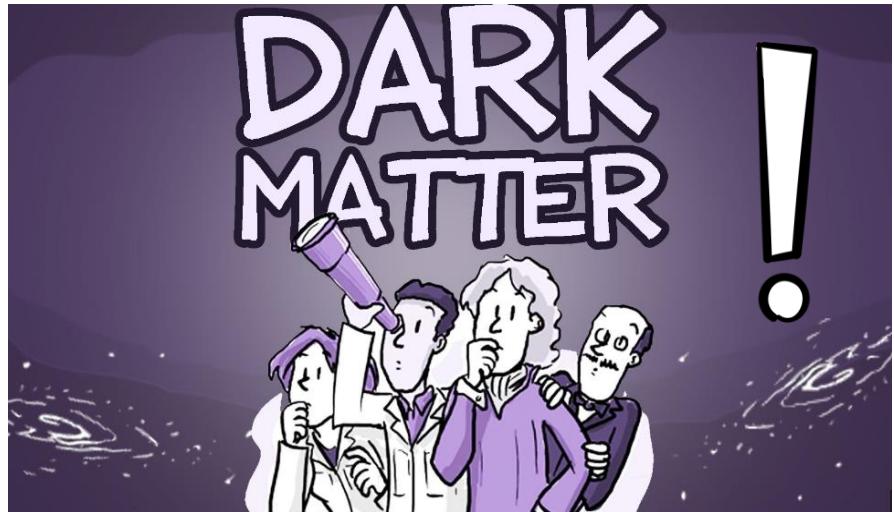


Q: J. Cham, CERN

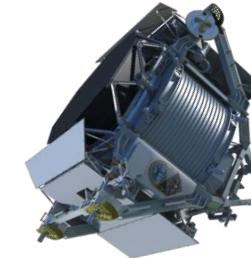
# Astro Particle Physics – open questions

## ■ what is the particle nature of Dark Matter?

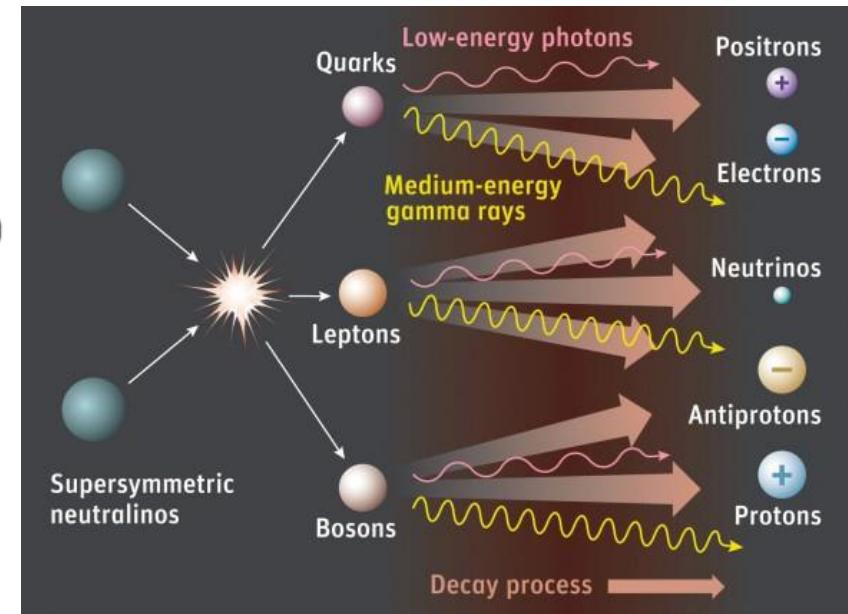
- we can find out by **observing its decay** in the cosmos!



galaxy



AMS\*\*



- using our knowledge of theory (SUSY)
- using our knowledge of cosmic accelerators

Q: J. Cham

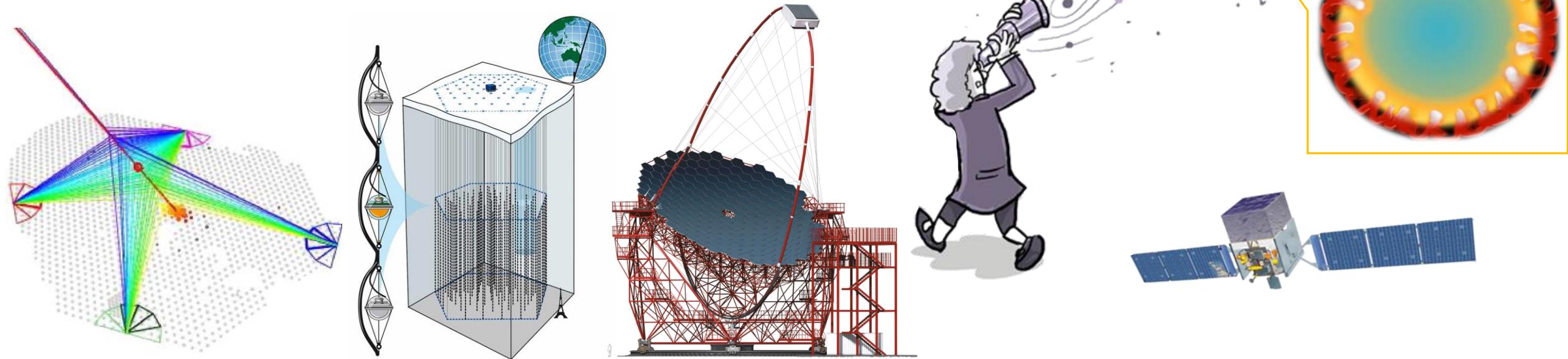
\* Cherenkov Telescope Array

\*\* Alpha Magnetic Spectrometer

# Astro Particle Physics – open questions

## ■ what is the nature of cosmic rays?

- we can find out by **using different messengers!**



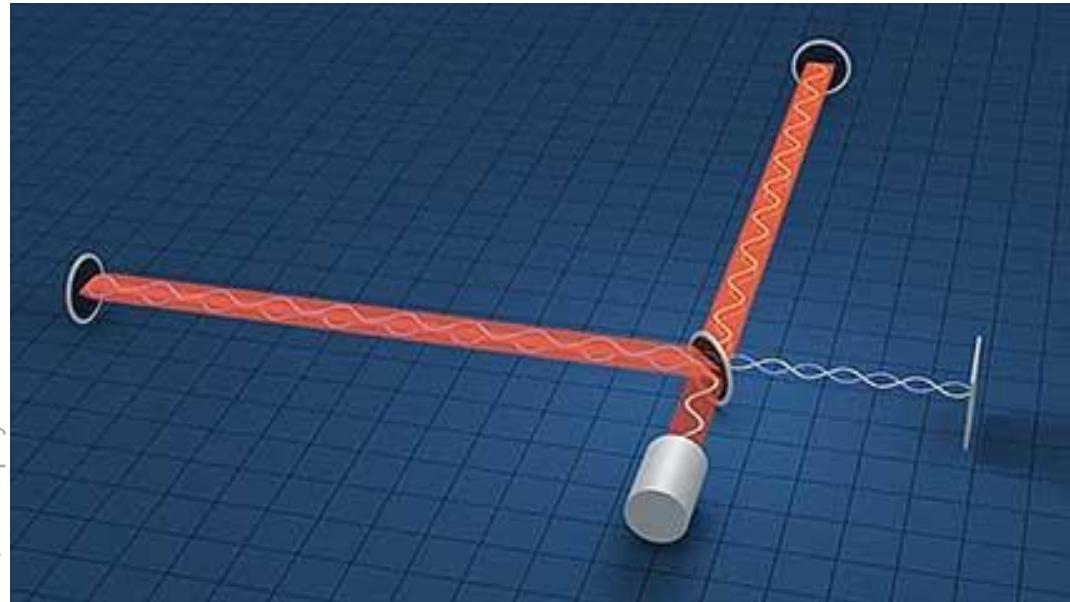
- combining our knowledge of each **messenger**: gammas, neutrinos, protons & nuclei
- using our knowledge of **magnetic fields**

Q: J. Cham

# Astro Particle Physics – open questions

## ■ what is the nature of gravitational waves

- we can find out by **using laser interferometers!**

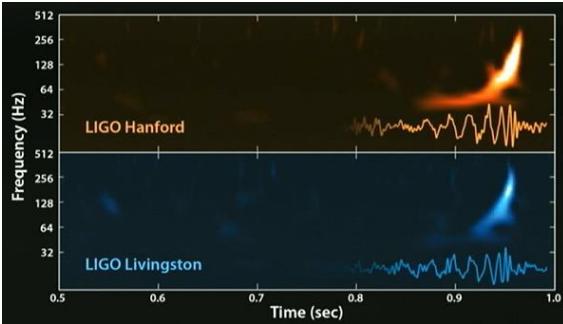


- using our knowledge of **general relativity**
- combining data from surface of earth & in space



# Astro Particle Physics – skills

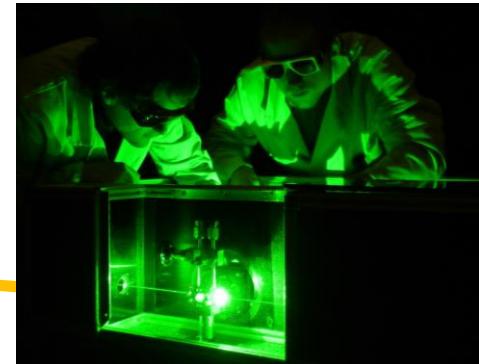
## ■ what skills do I learn in this lecture series?



**HOT TOPIC**

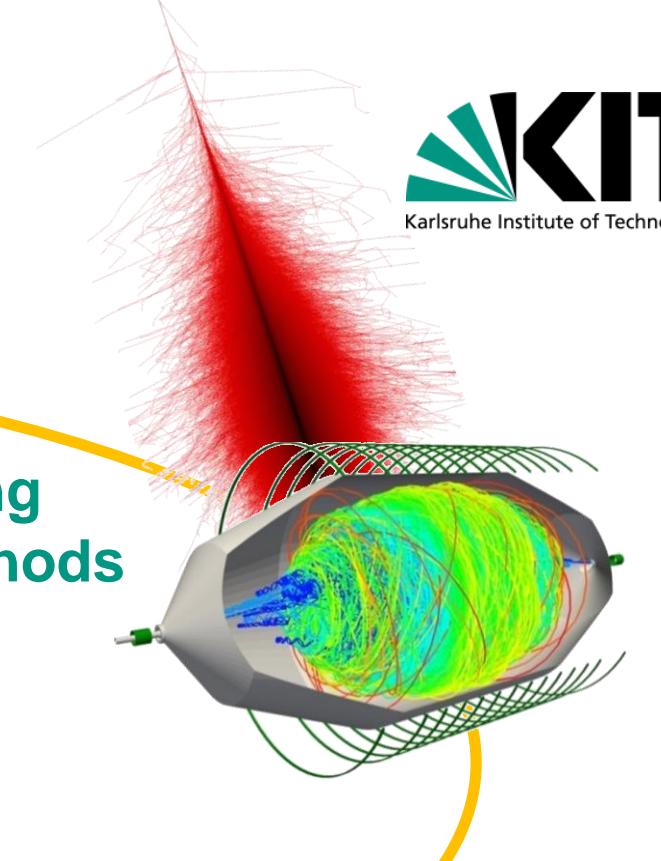


**hands-on**  
TRAINING



**modern detectors**

**advanced technologies**



**leading  
MC-methods**



**analysis techniques**

# Astro Particle Physics – skills

- what skills do I learn in this lecture series?



HOT TOPIC

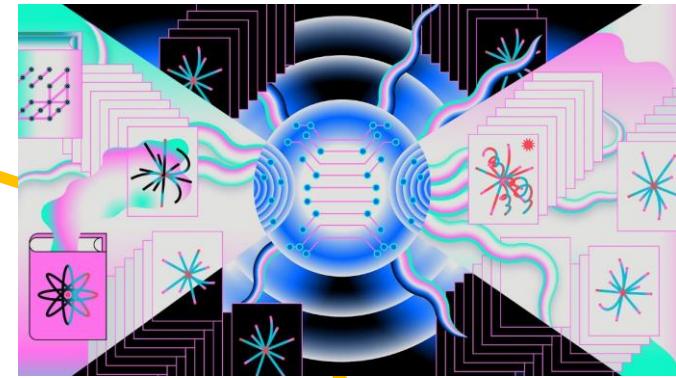


modern astrophysics



hands-on  
TRAINING

modern  
electronics



Q: symmetry magazine (2)



latest particle theory models

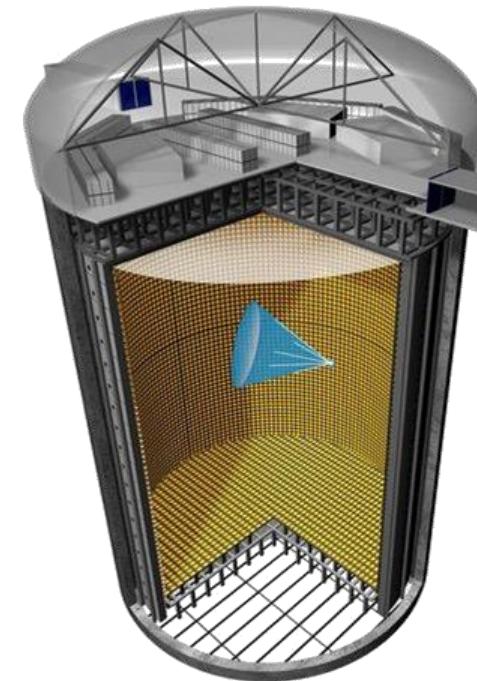
# Astro Particle Physics – science network

## ■ What connections does APP provide? Meet and greet... a Nobel Laureate

The Nobel Prize in Physics 2015 was awarded jointly to Takaaki Kajita and Arthur B. McDonald "for the discovery of neutrino oscillations, which shows that neutrinos have mass."



July 10, 2014 – Takaaki Kajita @ kl. HS A



**Nobelprize.org**  
The Official Web Site of the Nobel Prize

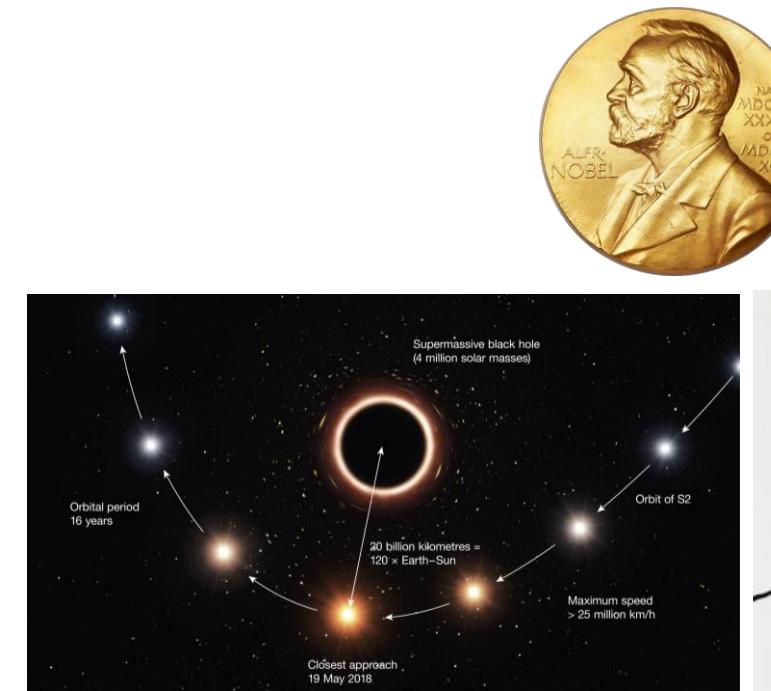


Takaaki Kajita

# Astro Particle Physics – science network

## ■ What connections does APP provide? Meet and greet... a Nobel Laureate

The Nobel Prize in Physics 2020 was divided, the other half jointly to Reinhard Genzel and Andrea Ghez "for the discovery of a supermassive compact object at the centre of our galaxy"



October 4/5, 2022 – R. Genzel @ KIT / KATRIN

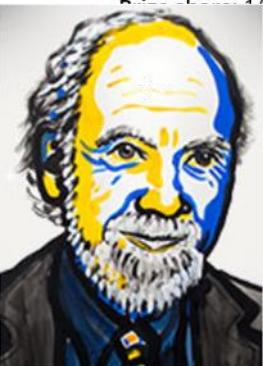
Reinhard Genzel

# Astro Particle Physics – recent breakthroughs

■ and the winner is...



© Nobel Media. Ill. N.  
Elmehed  
Rainer Weiss  
Prize share 1/2



© Nobel Media. Ill. N.  
Elmehed  
Barry C. Barish



© Nobel Media. Ill. N.  
Elmehed  
Kip S. Thorne



## Nobel Prize in Physics 2017

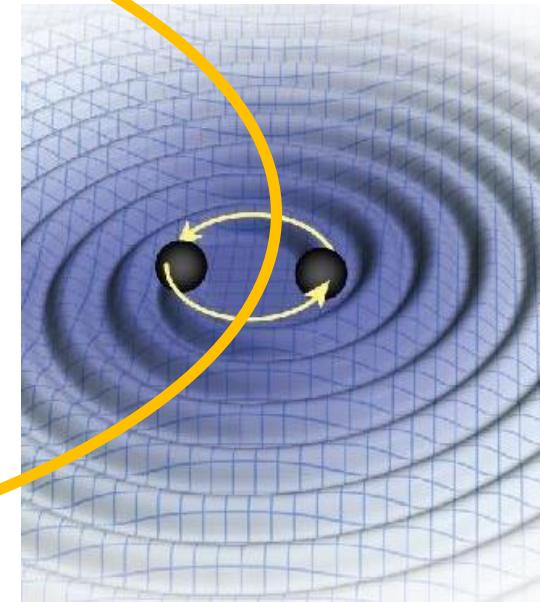
The Nobel Prize in Physics 2017 was divided, one half awarded to Rainer Weiss, the other half jointly to Barry C. Barish and Kip S. Thorne *"for decisive contributions to the LIGO detector and the observation of gravitational waves"*.

 **Nobelprize.org**

The Official Web Site of the Nobel Prize

Einstein's waves win Nobel Prize in physics

By Paul Rincon and Jonathan Amos  
BBC Science News



Q: BBC

# Astro Particle Physics – topic events (2021 ed.)

## ■ what was new in this lecture series? Interruption of TeV-gamma views...



direct view from MAGIC to  
volcanic eruption at the Cumbre  
Vieja (Santa Cruz de Tenerife)



# Astro Particle Physics – topic events (2022 ed.)



## ■ what is new this semester in this lecture series?

First Dark Matter Search Results from the LUX-ZEPLIN (LZ) Experiment

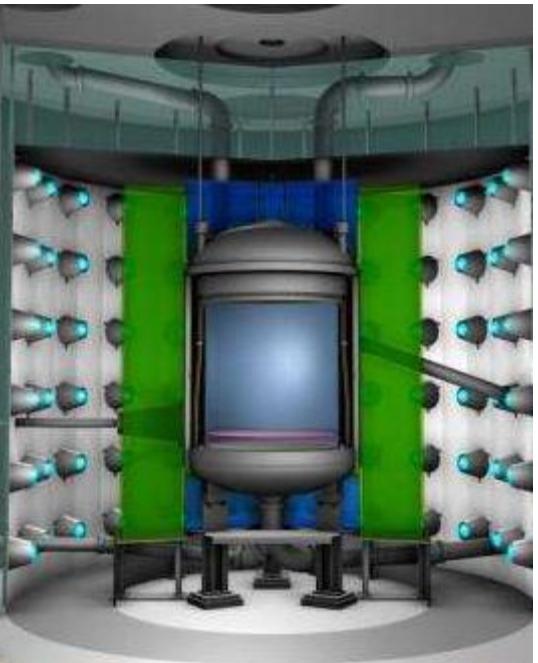
J. Aalbers,<sup>1,2</sup> D.S. Akerib,<sup>1,2</sup> C.W. Akerlof,<sup>3</sup> A.K. Al Musalhi,<sup>4</sup> F. Alder,<sup>5</sup> A. Alqahtani,<sup>6</sup> S.K. Alsum,<sup>7</sup> C.S. Amarasinghe,<sup>3</sup> A. Ames,<sup>1,2</sup> T.J. Anderson,<sup>1,2</sup> N. Angelides,<sup>5,8</sup> H.M. Araujo,<sup>8</sup> J.E. Armstrong,<sup>9</sup> M. Arthurs,<sup>3</sup> S. Azadi,<sup>10</sup> A.J. Bailey,<sup>8</sup> A. Baker,<sup>8</sup> J. Balajthy,<sup>11</sup> S. Balashov,<sup>12</sup> J. Bang,<sup>6</sup> J.W. Bargemann,<sup>10</sup> M.J. Barry,<sup>13</sup> J. Barthel,<sup>14</sup> D. Bauer,<sup>8</sup> A. Baxter,<sup>15</sup> K. Beattie,<sup>13</sup> J. Belle,<sup>16</sup> P. Beltrame,<sup>5,17</sup> J. Bensinger,<sup>18</sup> T. Benson,<sup>7</sup> E.P. Bernard,<sup>13,19</sup> A. Bhatti,<sup>9</sup> A. Biekert,<sup>13,19</sup> T.P. Biesiadzinski,<sup>1,2</sup> H.J. Birch,<sup>3,15</sup> B. Birrittella,<sup>7</sup> G.M. Blockinger,<sup>20</sup> K.E. Boat,<sup>4</sup> B. Boxer,<sup>11,15</sup> R. Bramante,<sup>1,2</sup> C.A.J. Brew,<sup>12</sup> P. Brás,<sup>21</sup> J.H. Buckley,<sup>22</sup> V.V. Bugaev,<sup>22</sup> S. Burdin,<sup>15</sup> J.K. Busenitz,<sup>23</sup> M. Buuck,<sup>1,2</sup> R. Cabrita,<sup>21</sup> C. Carels,<sup>4</sup> D.L. Carlsmith,<sup>7</sup> B. Carlson,<sup>14</sup> M.C. Carmona-Benitez,<sup>24</sup> M. Casella,<sup>5</sup> C. Chan,<sup>6</sup> A. Chawla,<sup>25</sup> H. Chen,<sup>13</sup> J.J. Cherwinka,<sup>7</sup> N.I. Chott,<sup>26</sup> A. Cole,<sup>13</sup> J. Coleman,<sup>13</sup> M.V. Converse,<sup>27</sup> A. Cottle,<sup>4,16</sup> G. Cox,<sup>14,24</sup> W.W. Craddock,<sup>1</sup> O. Creaner,<sup>13</sup> D. Curran,<sup>14</sup> A. Currie,<sup>8</sup> J.E. Cutter,<sup>11</sup> C.E. Dahl,<sup>16,28</sup> A. David,<sup>5</sup> J. Davis,<sup>14</sup> T.J.R. Davison,<sup>17</sup> J. Delgaudio,<sup>14</sup> S. Dey,<sup>4</sup> L. de Viveiros,<sup>24</sup> A. Dobi,<sup>13</sup> J.E.Y. Dobson,<sup>5</sup> E. Druszkiewicz,<sup>27</sup> A. Dushkin,<sup>18</sup> T.K. Edberg,<sup>9</sup> W.R. Edwards,<sup>13</sup> M.M. Elmim,<sup>23</sup> W.T. Emmet,<sup>29</sup> S.R. Erikson,<sup>30</sup> C.H. Faham,<sup>13</sup> A. Fan,<sup>1,2,\*</sup> S. Fayer,<sup>8</sup> N.M. Fearnor,<sup>4</sup> S. Fiorucci,<sup>13</sup> H. Flaeche,<sup>30</sup> P. Ford,<sup>12</sup> V.B. Francis,<sup>12</sup> E.D. Frasier,<sup>15</sup> T. Fruth,<sup>4,5</sup> R.J. Gaitskell,<sup>6</sup> N.J. Gantos,<sup>13</sup> D. Garcia,<sup>6</sup> A. Geffre,<sup>14</sup> V.M. Gehman,<sup>13</sup> J. Genovesi,<sup>26</sup> C. Ghag,<sup>5</sup> R. Gibbons,<sup>13,19</sup> E. Gibson,<sup>4</sup> M.G.D. Gilchriese,<sup>13</sup> S. Gokhale,<sup>31</sup> B. Gomber,<sup>7</sup> J. Green,<sup>4</sup> A. Greenall,<sup>15</sup> S. Greenwood,<sup>8</sup> M.G.D.van der Grinten,<sup>12</sup> C.B. Gwilliam,<sup>15</sup> C.R. Hall,<sup>9</sup> S. Hans,<sup>31</sup> K. Hanzel,<sup>13</sup> A. Harrison,<sup>26</sup> E. Hartigan-O'Connor,<sup>6</sup> S.J. Haselschwandt,<sup>13</sup> S.A. Hertel,<sup>32</sup> G. Heuermann,<sup>3</sup> C. Hjemfelt,<sup>26</sup> M.D. Hoff,<sup>13</sup> E. Holtom,<sup>12</sup> J.Y.-K. Hor,<sup>23</sup> M. Horn,<sup>14</sup> D.Q. Huang,<sup>3,6</sup> D. Hunt,<sup>4</sup> C.M. Ignarra,<sup>1,2</sup> R.G. Jacobsen,<sup>13,19</sup> O. Jahangir,<sup>5</sup> R.S. James,<sup>5</sup> S.N. Jeffery,<sup>12</sup> W. Ji,<sup>1,2</sup> J. Johnson,<sup>11</sup> A.C. Kaboth,<sup>12,25,†</sup> A.C. Kamaha,<sup>20,33</sup> V. Kamenid,<sup>13,19</sup> V. Kasey,<sup>8</sup> K. Kazkaz,<sup>34</sup> J. Keefer,<sup>14</sup> D. Khaitan,<sup>27</sup> M. Khaleeq,<sup>8</sup> A. Khazov,<sup>12</sup> I. Khurana,<sup>5</sup> Y.D. Kim,<sup>35</sup> C.D. Kocher,<sup>6</sup> D. Kodroff,<sup>24</sup> L. Korley,<sup>3,18</sup> E.V. Korolkova,<sup>36</sup> J. Kras,<sup>7</sup> H. Kraus,<sup>4</sup> S. Kravitz,<sup>13</sup> H.J. Krebs,<sup>1</sup> L. Kreczko,<sup>30</sup> B. Krikler,<sup>30</sup> V.A. Kudryavtsev,<sup>36</sup> S. Kyre,<sup>10</sup> B. Landerud,<sup>7</sup> E.A. Leason,<sup>17</sup> C. Lee,<sup>1,2</sup> J. Lee,<sup>35</sup> D.S. Leonard,<sup>35</sup> R. Leonard,<sup>26</sup> K.T. Lesko,<sup>13</sup> C. Levy,<sup>20</sup> J. Li,<sup>35</sup> F.-T. Liao,<sup>4</sup> J. Liao,<sup>6</sup> J. Lin,<sup>4,13,19</sup> A. Lindote,<sup>21</sup> R. Linehan,<sup>1,2</sup> W.H. Lippincott,<sup>10,16</sup> R. Liu,<sup>6</sup> X. Liu,<sup>17</sup> Y. Liu,<sup>7</sup> C. Loniewski,<sup>27</sup> M.I. Lopes,<sup>21</sup> E. Lopez Asamar,<sup>21</sup> B. López Perez-Paredes,<sup>8</sup> W. Lorenzon,<sup>3</sup> D. Lucero,<sup>14</sup> S. Luitz,<sup>1</sup> J.M. Lyle,<sup>6</sup> P.A. Majewski,<sup>12</sup> J. Makkinje,<sup>6</sup> D.C. Malling,<sup>6</sup> A. Manalaysay,<sup>11,13</sup> L. Manenti,<sup>5</sup> R.L. Mammino,<sup>7</sup> N. Marangou,<sup>8</sup> M.F. Marzioni,<sup>17</sup> C. Maupin,<sup>14</sup> M.E. McCarthy,<sup>27</sup> C.T. McConnell,<sup>13</sup> D.N. McKinsey,<sup>13,19</sup> J. McLaughlin,<sup>28</sup> Y. Meng,<sup>23</sup> J. Mignault,<sup>6</sup> E.H. Miller,<sup>1,2,26</sup> E. Mizrachi,<sup>9,34</sup> J.A. Mock,<sup>13,20</sup> A. Monte,<sup>10,16</sup> M.E. Monzani,<sup>1,2,37</sup> J.A. Morad,<sup>11</sup> J.D. Morales Mendoza,<sup>1,2</sup> E. Morrison,<sup>26</sup> B.J. Mount,<sup>38</sup> M. Murdy,<sup>32</sup> A.St.J. Murphy,<sup>17</sup> D. Nain,<sup>11</sup> A. Naylor,<sup>26</sup> C. Nedlik,<sup>32</sup> C. Nehrkorn,<sup>10</sup> H.N. Nelson,<sup>10</sup> F. Neves,<sup>21</sup> A. Nguyen,<sup>17</sup> J.A. Nikoleyeczik,<sup>7</sup> A. Nilima,<sup>17</sup> J. O'Dell,<sup>12</sup> F.G. O'Neill,<sup>1</sup> K. O'Sullivan,<sup>13,19</sup> I. Olcina,<sup>13,19</sup> M.A. Olevitch,<sup>22</sup> K.C. Oliver-Mallory,<sup>8,13,19</sup> J. Orpwood,<sup>36</sup> D. Pagenkopf,<sup>10</sup> S. Pal,<sup>21</sup> K.J. Palladino,<sup>4,7</sup> J. Palmer,<sup>25</sup> M. Pangilinan,<sup>6</sup> N. Parveen,<sup>20</sup> S.J. Patton,<sup>13</sup> E.K. Pease,<sup>13</sup> B. Penning,<sup>3,18</sup> C. Pereira,<sup>21</sup> G. Pereira,<sup>5</sup> T. Persingh,<sup>34</sup> I.B. Peterson,<sup>13</sup> A. Piepke,<sup>23</sup> J. Podzerwinski,<sup>7</sup> D. Porzio,<sup>21,‡</sup> S. Powell,<sup>15</sup> R.M. Preece,<sup>12</sup> K. Pushkin,<sup>3</sup> Y. Qie,<sup>27</sup> B.N. Ratcliff,<sup>1</sup> J. Reichenbacher,<sup>26</sup> L. Reichhart,<sup>5</sup> C.A. Rhyne,<sup>6</sup> A. Richards,<sup>8</sup> Q. Riffard,<sup>13,19</sup> G.R.C. Rischbieter,<sup>20</sup> J.P. Rodrigues,<sup>21</sup> A. Rodriguez,<sup>38</sup> H.J. Rose,<sup>15</sup> R. Rosero,<sup>31</sup> P. Rossiter,<sup>36</sup> T. Rushton,<sup>36</sup> G. Rutherford,<sup>6</sup> D. Rynders,<sup>14</sup> J.S. Saba,<sup>13</sup> D. Santone,<sup>25</sup> A.B.M.R. Sazzad,<sup>23</sup> R.W. Schnee,<sup>26</sup> P.R. Scovell,<sup>4,12</sup> D. Seymour,<sup>6</sup> S. Shaw,<sup>10</sup> T. Shutt,<sup>1,2</sup> J.J. Silk,<sup>9</sup> C. Silva,<sup>21</sup> G. Sinev,<sup>26</sup> K. Skarpas,<sup>1</sup> W. Skulski,<sup>27</sup> R. Smith,<sup>13,19</sup> M. Solmaz,<sup>10</sup> V.N. Solovov,<sup>21</sup> P. Sorensen,<sup>13</sup> J. Soria,<sup>13,19</sup> I. Stancu,<sup>23</sup> M.R. Stark,<sup>26</sup> A. Stevens,<sup>4,5,8</sup> T.M. Stiegler,<sup>39</sup> K. Stifter,<sup>1,2,16</sup> R. Studley,<sup>18</sup> B. Suerfu,<sup>13,19</sup> T.J. Summer,<sup>8</sup> P. Sutcliffe,<sup>15</sup> N. Swanson,<sup>6</sup> M. Szydagis,<sup>20</sup> M. Tan,<sup>4</sup> D.J. Taylor,<sup>14</sup> R. Taylor,<sup>8</sup> W.C. Taylor,<sup>6</sup> D.J. Temples,<sup>28</sup> B.P. Tenmyer,<sup>29</sup> P.A. Terman,<sup>39</sup> K.J. Thomas,<sup>13</sup> D.R. Tiedt,<sup>9,14,26</sup> M. Timalsina,<sup>26</sup> W.H. To,<sup>1,2</sup> A. Tomás,<sup>8</sup> Z. Tong,<sup>8</sup> D.R. Tovey,<sup>36</sup> J. Tranter,<sup>36</sup> M. Trask,<sup>10</sup> M. Tripathi,<sup>11</sup> D.R. Troutstad,<sup>26</sup> C.E. Tull,<sup>13</sup> W. Turner,<sup>15</sup> L. Tvrznikova,<sup>19,29,34</sup> U. Utiku,<sup>5</sup> J. Va'vra,<sup>1</sup> A. Vacheret,<sup>8</sup> A.C. Vaitkus,<sup>6</sup> J.R. Verbus,<sup>6</sup> E. Voisin,<sup>16</sup> W.L. Waldron,<sup>13</sup> A. Wang,<sup>1,2</sup> B. Wang,<sup>23</sup> J.J. Wang,<sup>23</sup> W. Wang,<sup>7,32</sup> Y. Wang,<sup>13,19</sup> J.R. Watson,<sup>13,19</sup> R.C. Webb,<sup>39</sup> A. White,<sup>6</sup> D.T. White,<sup>10</sup> J.T. White,<sup>39,4</sup> R.G. White,<sup>1,2</sup> T.J. Whitis,<sup>1,10</sup> M. Williams,<sup>3,18</sup> W.J. Wisniewski,<sup>1</sup> M.S. Witherell,<sup>13,19</sup> F.L.H. Wolfs,<sup>27</sup> J.D. Wolfs,<sup>27</sup> S. Woodford,<sup>15</sup> D. Woodward,<sup>24,§</sup> S.D. Worm,<sup>12</sup> C.J. Wright,<sup>30</sup> Q. Xia,<sup>13</sup> X. Xiang,<sup>6,31</sup> Q. Xiao,<sup>7</sup> J. Xu,<sup>34</sup> M. Yeh,<sup>31</sup> J. Yin,<sup>27</sup> I. Young,<sup>16</sup> P. Zarzhitsky,<sup>23</sup> A. Zuckerman,<sup>6</sup> and E.A. Zweig,<sup>33</sup> (The LUX-ZEPLIN (LZ) Collaboration)

arXiv:2207.03764v2 [hep-ex] 18 Jul 2022



2207.03764.pdf (arxiv.org)

New results on  
WIMP Dark Matter



Exp. Particle Physics - ETP

World's most sensitive  
dark matter detector  
tested for the first time

A brief test has proven that the new LUX-ZEPLIN dark matter detector is the most sensitive ever. It may be our best bet for finally finding dark matter particles

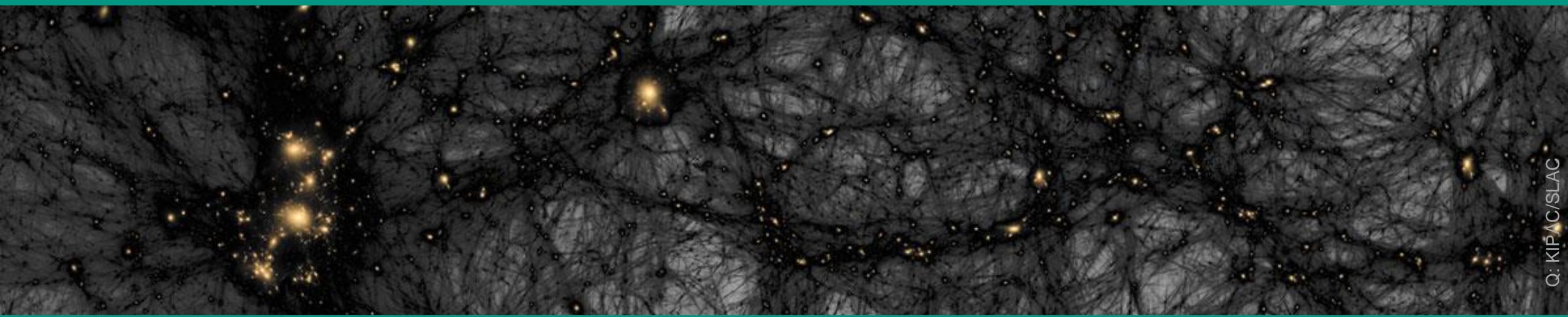


PHYSICS 7 July 2022

By Leah Crane



The LUX-ZEPLIN dark matter detector at the Sanford Underground Research Facility in South Dakota is the most sensitive ever



Q: KIPAC/SLAC

# OUTLINE

## 1. Introduction

- 1.1      **Particle Radiation from the Laboratory**
- 1.2      **Particle Radiation from the Universe**

## 2. Experimental Methods

- 2.1      **Multi-Messenger Methods**
  - 2.1.1      Air Shower Experiments
  - 2.1.2      Gamma Telescopes
  - 2.1.3      Neutrino Telescopes

## 2.2      **Search for Rare Events**

2.2.1      Background processes

2.2.2      Shielding methods

2.2.3      Primordial decay chains

## 3.      **Neutrinos**

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3.1      Introduction

3.2      **kinematic determination of the neutrino mass**

3.3      **Search for  $0\nu\beta\beta$  processes**

## 4. Dark Matter

- 4.1 Introduction
- 3.2 DM Candidates
- 3.3 WIMP search at the LHC
- 3.4 indirect WIMP detection methods
  - 3.4.1 Gammas and positrons
  - 3.4.2 Neutrinos

# Astroparticle Physics – 1: Outline

## 4.5 direct detection methods for WIMPs

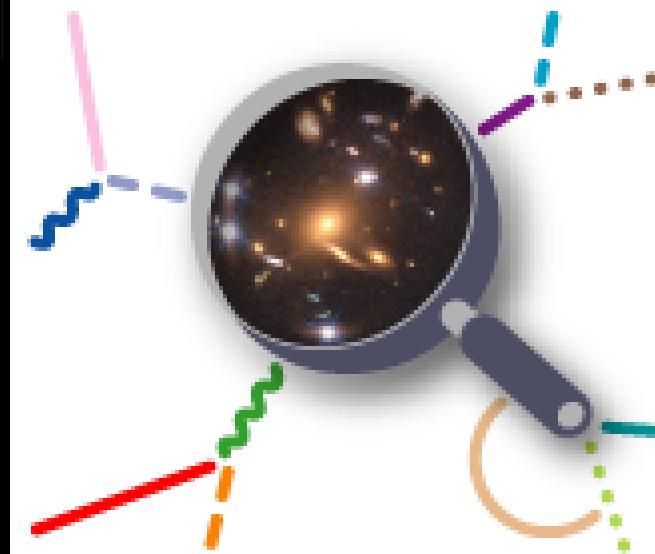
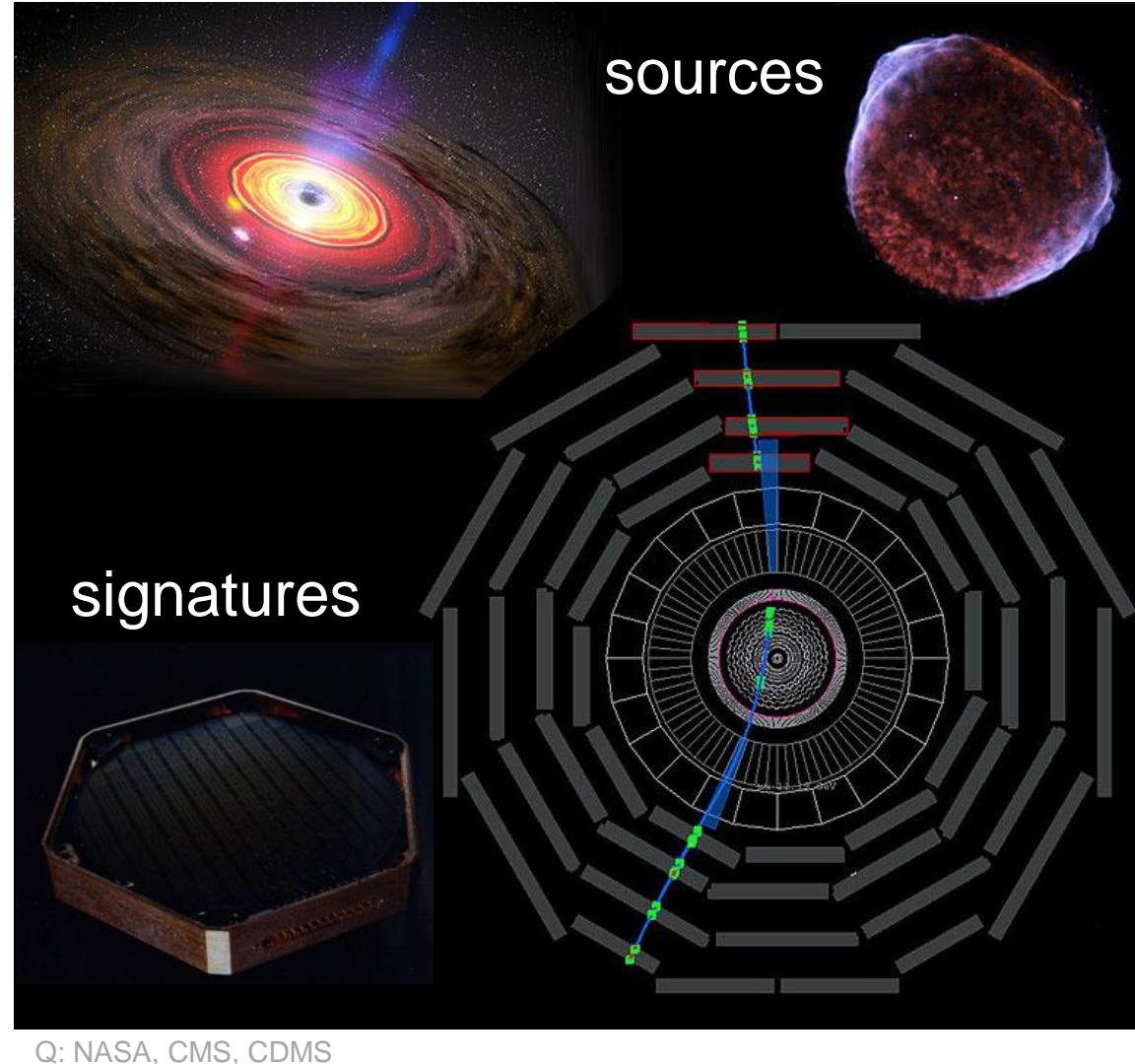
- 4.5.1 reaction kinematics
- 4.5.2 cryogenic bolometers
- 4.5.3 liquid noble gas detectors
- 4.5.4 future projects & outlook

## 4.6 non-thermal DM candidates

- 4.6.1 axions
- 4.6.2 keV neutrinos

# 1. INTRODUCTION

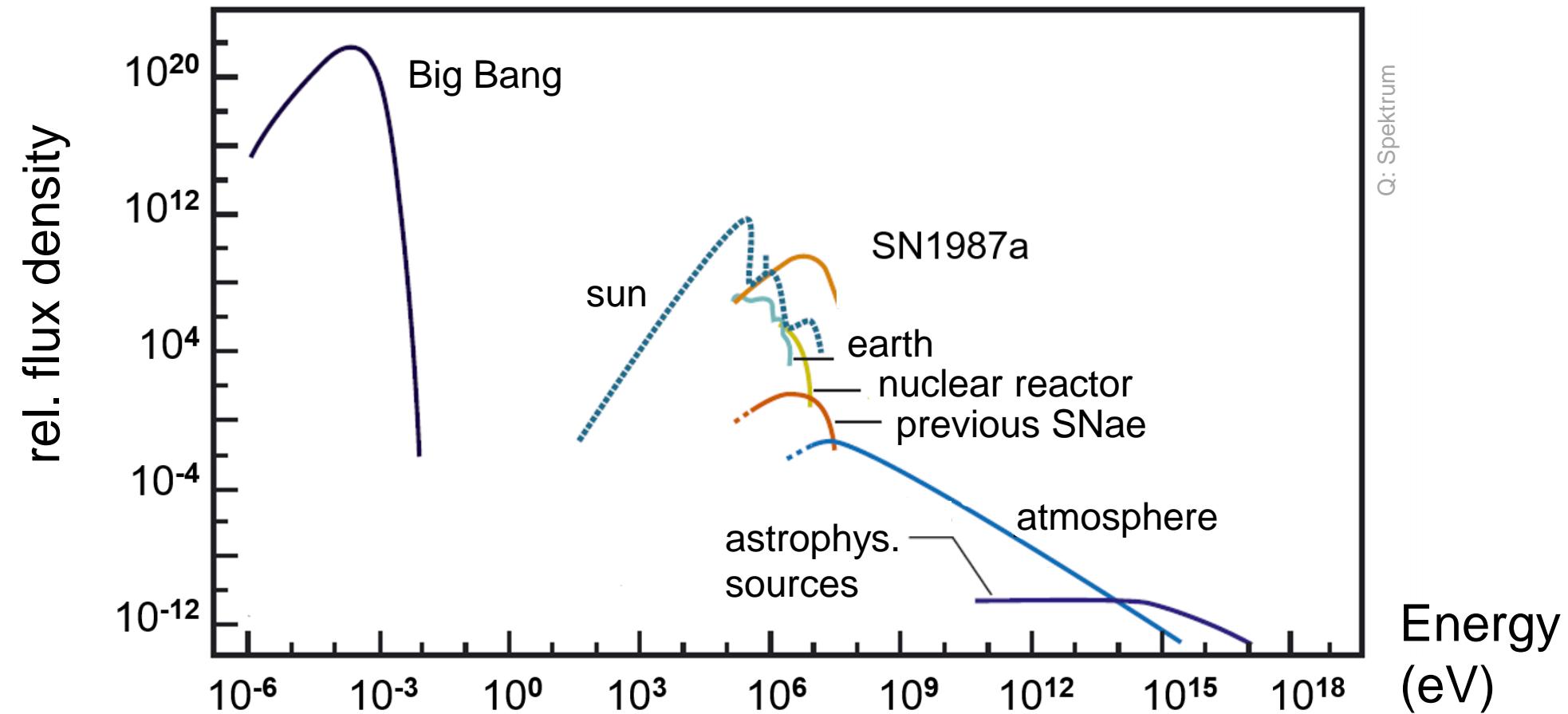
## ■ Particle radiation from the **Laboratory** and from the Universe



detection methods

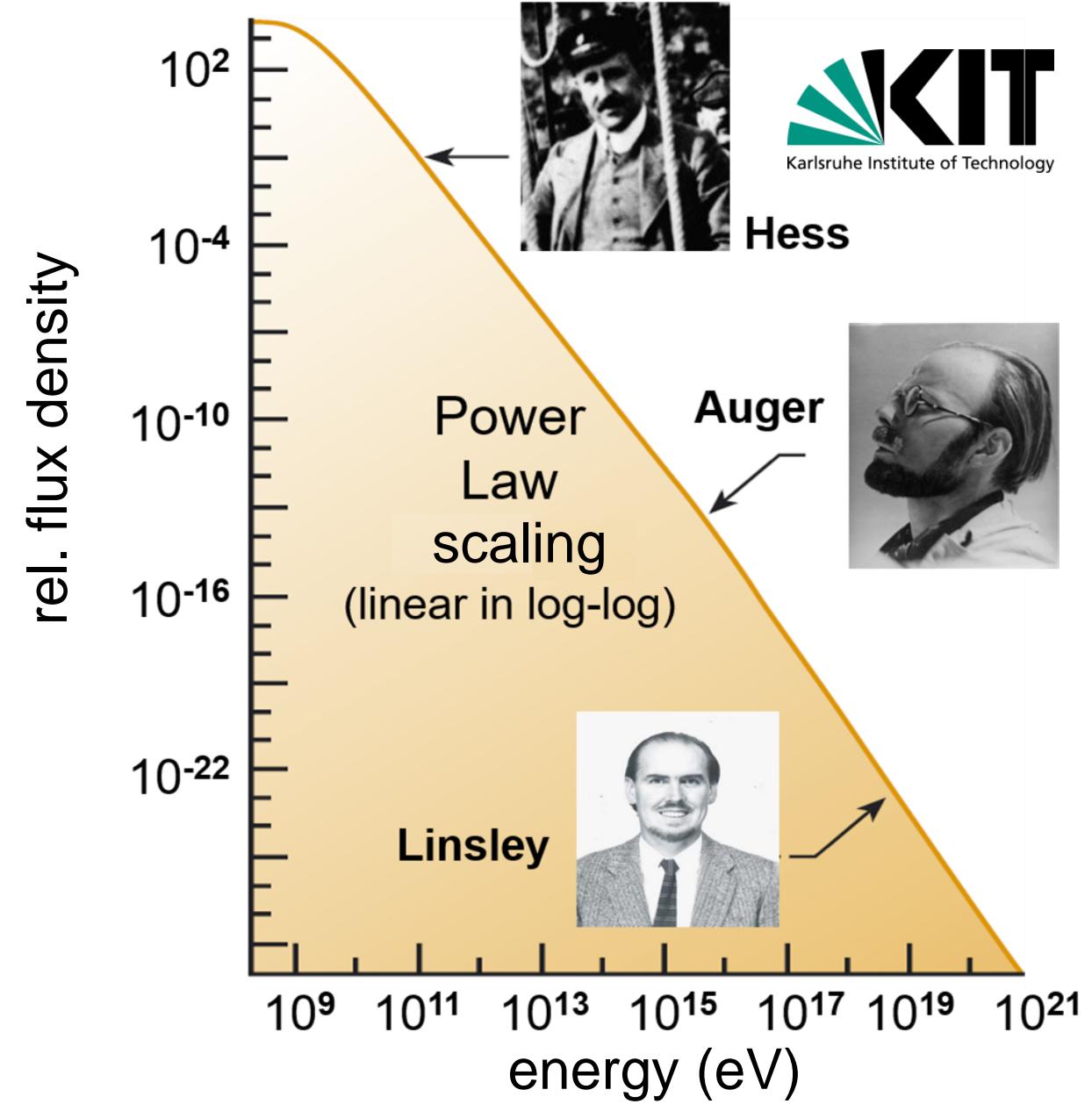
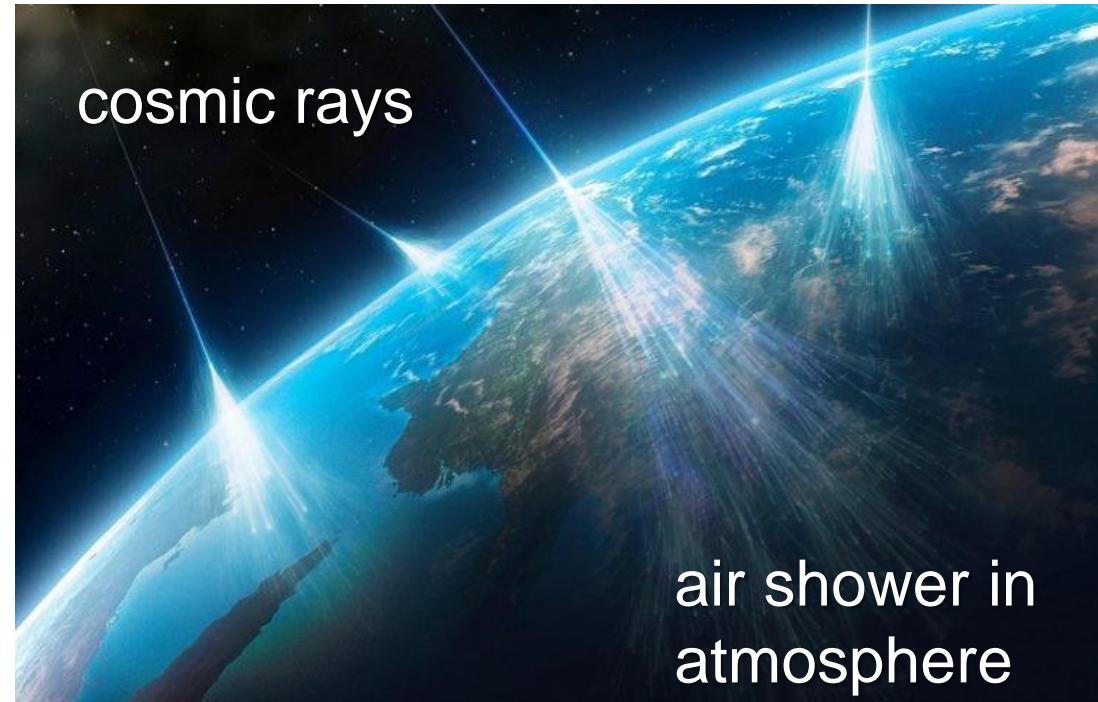
# Example: neutrinos from the earth & beyond

- extremely broad energy region, many sources



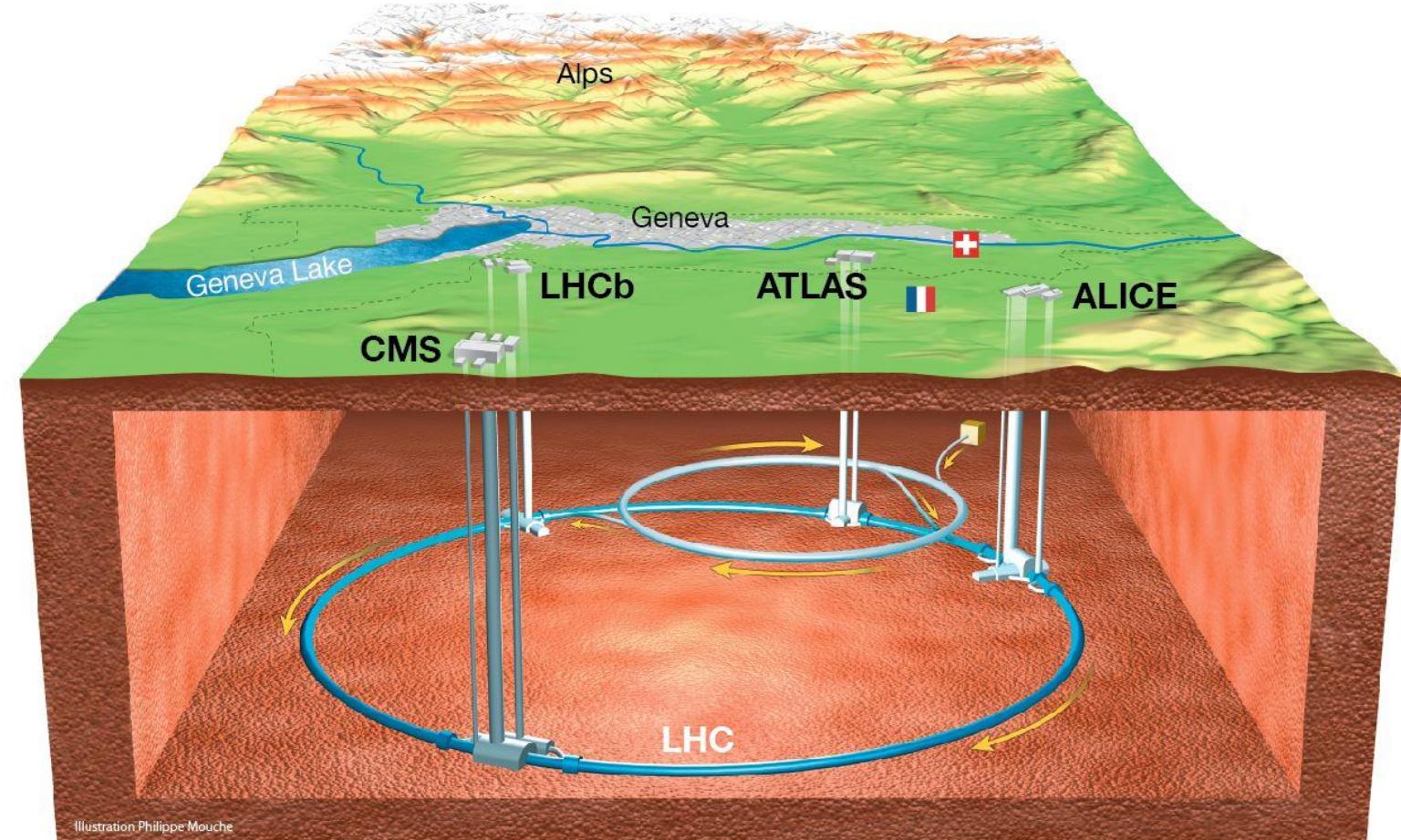
# Example: cosmic rays

- extremely broad energy region,  
many sources

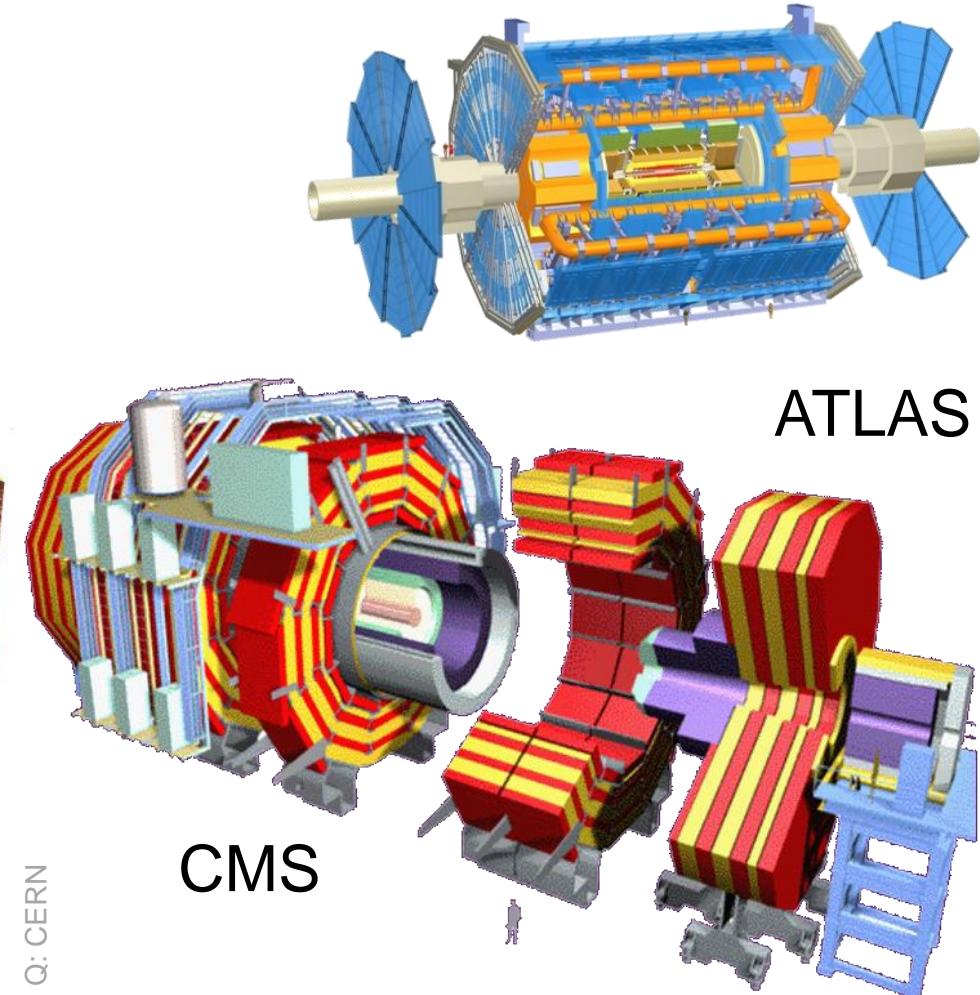


# 1.1 particle radiation from the laboratory

## ■ accelerators at TeV-scale: Search for SUSY\*

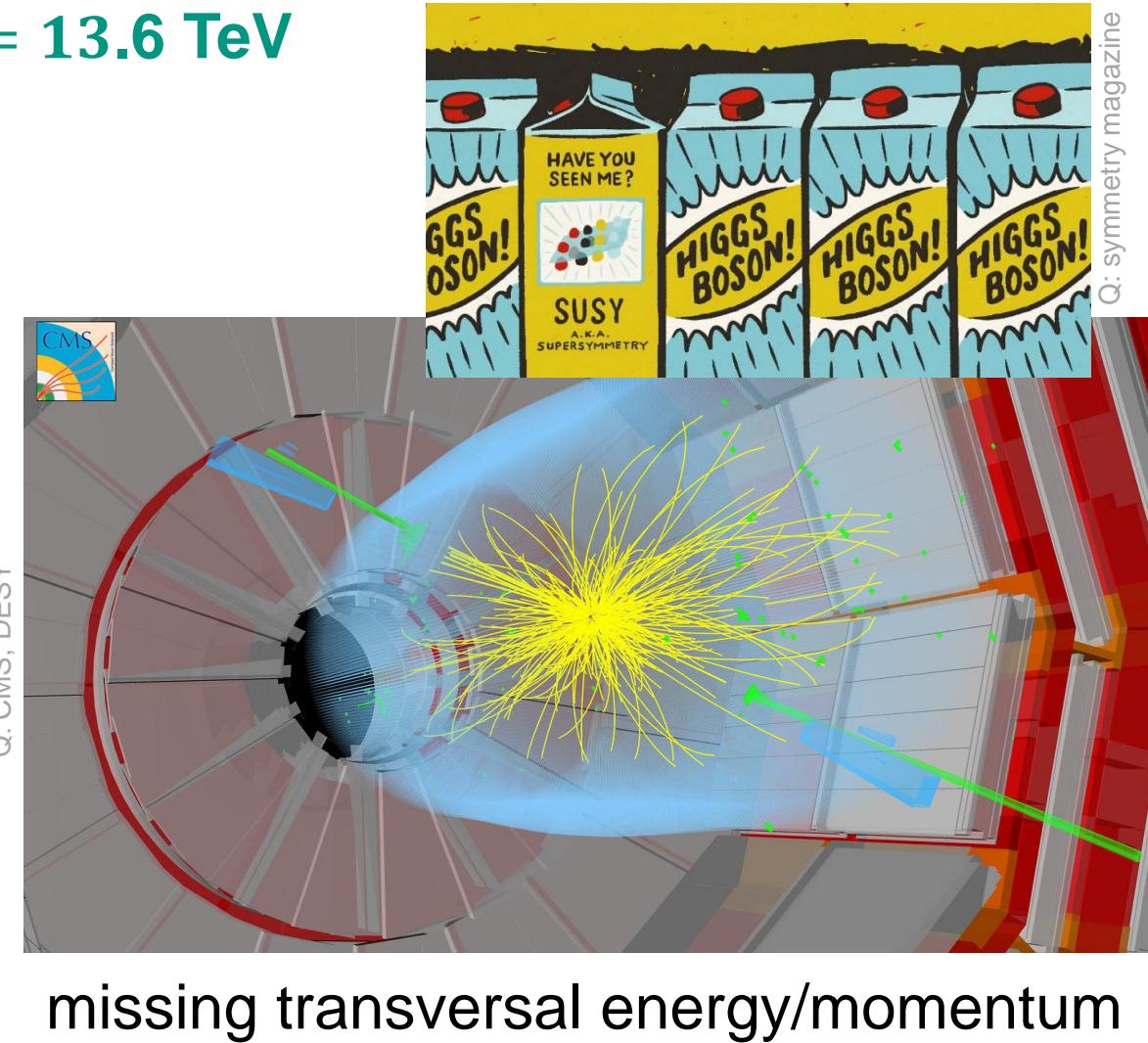
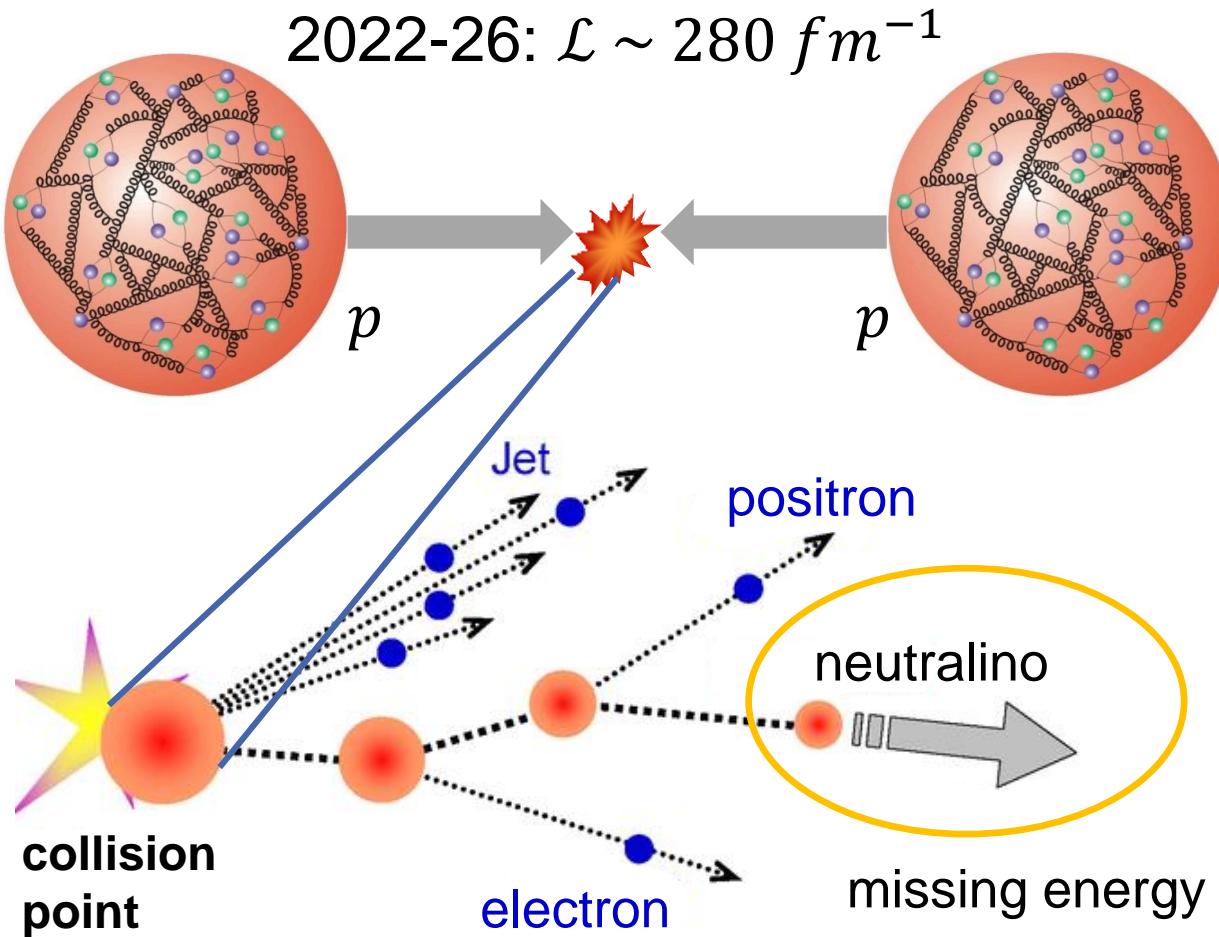


Q: CERN



# Accelerator-based SUSY-searches at LHC

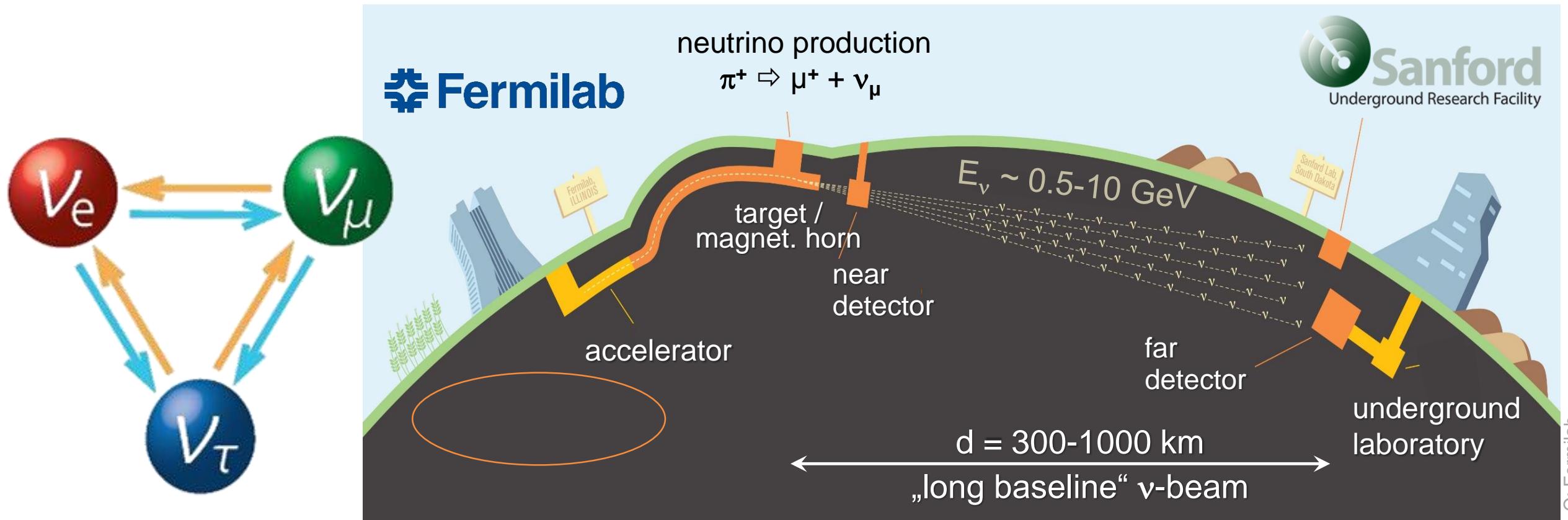
## ■ Run 3: proton-proton collisions at $\sqrt{s} = 13.6 \text{ TeV}$



# Accelerator-based neutrino oscillation studies

- long-baseline (LBL) neutrino beams at the **GeV-scale**: mixing properties, CP

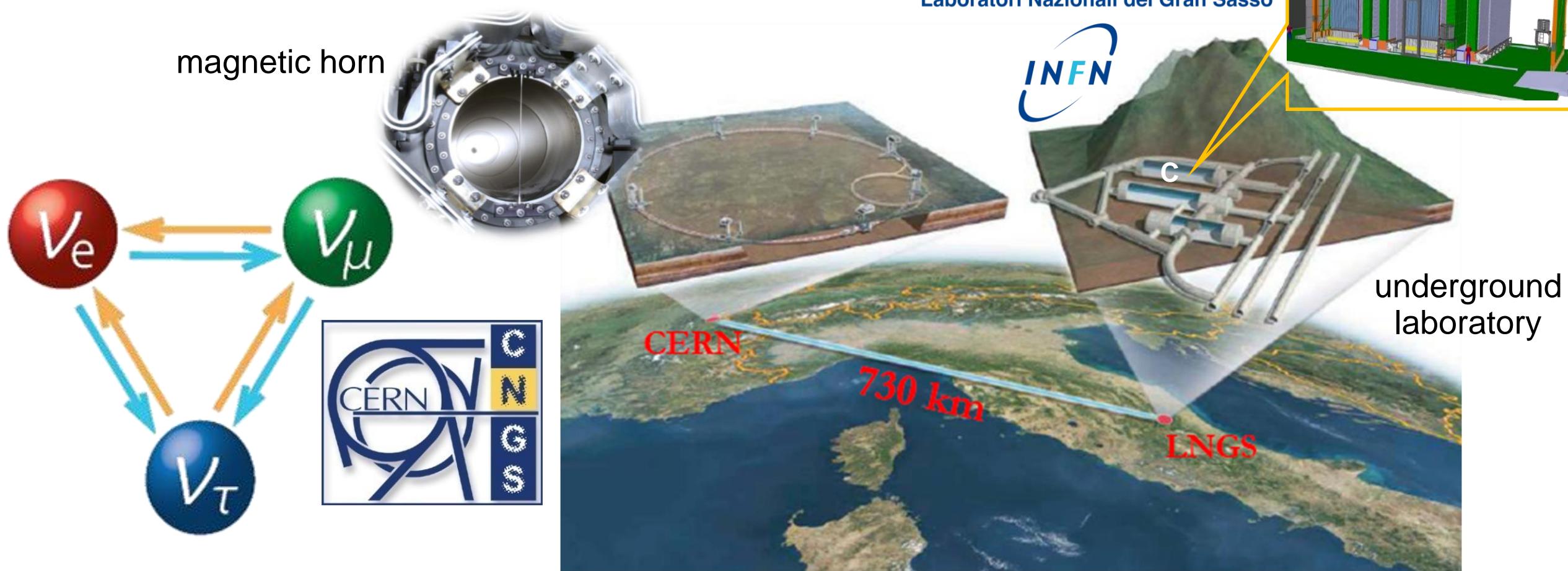
set-up of an LBL- $\nu$ -experiment



Q: Fermilab

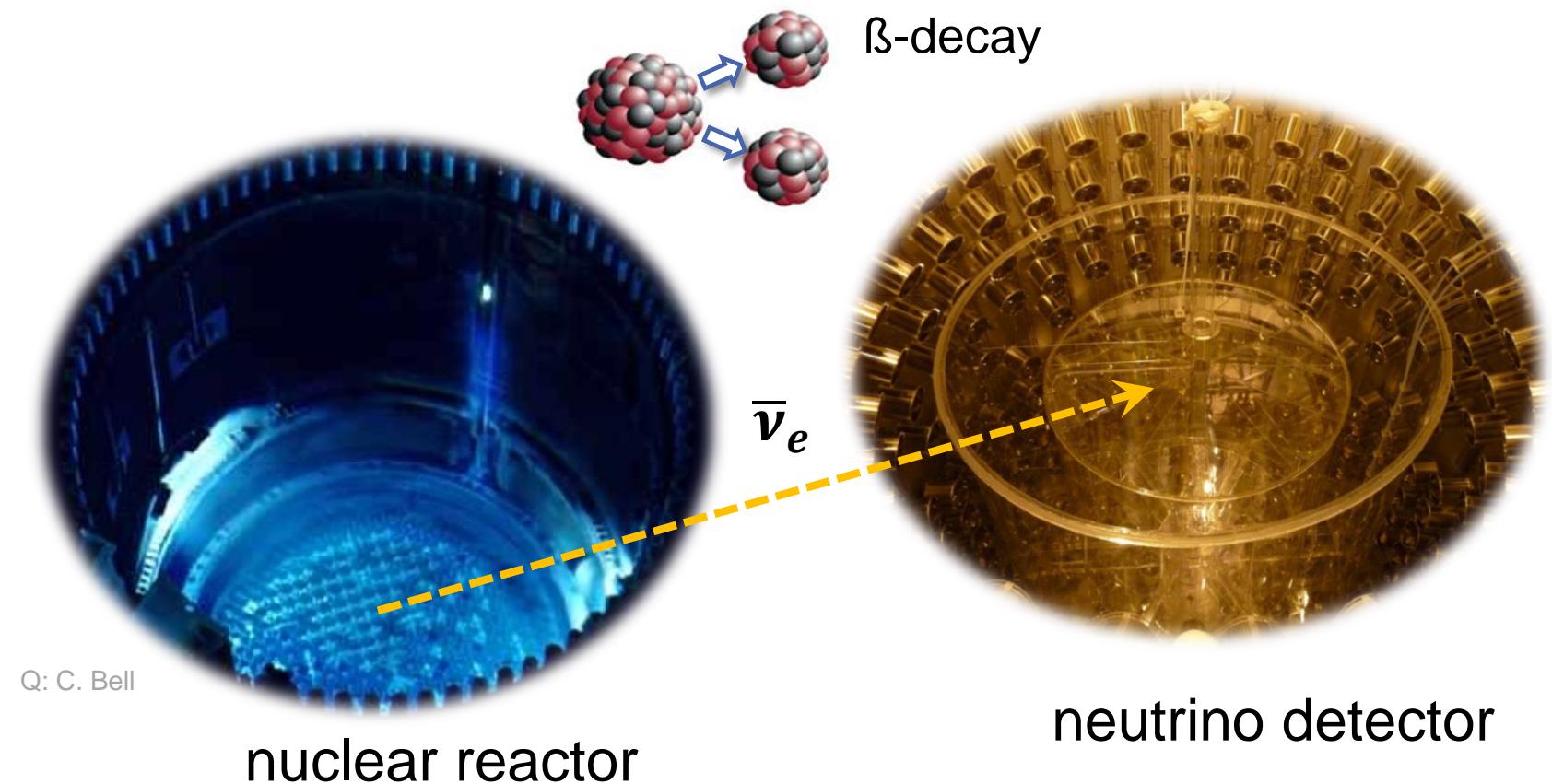
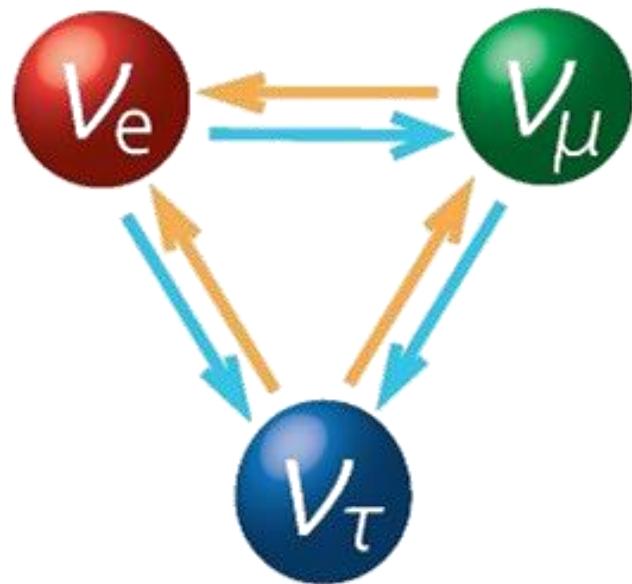
# Accelerator-based neutrino oscillation studies

## ■ example: the CERN – Gran Sasso LBL-neutrino-beam



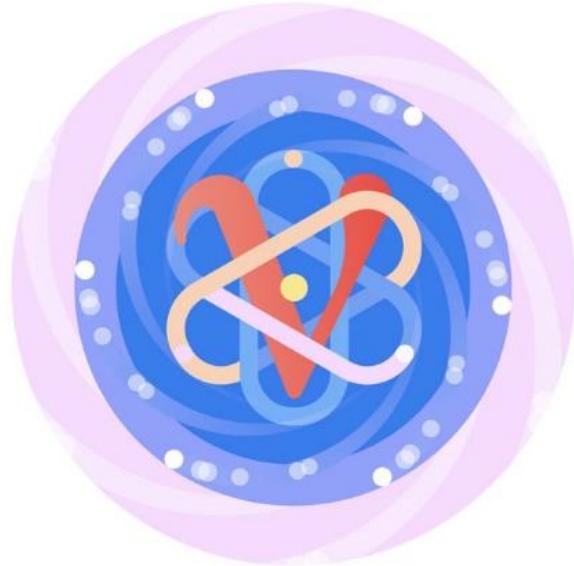
# reactor-based neutrino oscillation studies

- long-baseline neutrino beams at the **MeV-scale**: mixing properties, CP



# reactor-based neutrino oscillation studies

## ■ example: the Daya Bay neutrino-oscillation experiment in China



Q: Daya Bay

