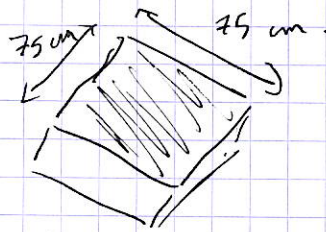


# COSMIC RAYS - Exercises set 1

## 1. Cosmic rays at Earth



$$\rightarrow \begin{cases} A = 0.5625 \text{ m}^2 \\ \Omega = A \pi \sin^2 \theta_{\max} \times 2? \\ \theta_{\max} = 80^\circ \Rightarrow \Omega = 1.7 \text{ sr} \end{cases}$$

$$\phi(\epsilon) = \frac{d^4 N}{dE dA d\Omega dt} = \phi_1 \left( \frac{\epsilon}{E_1} \right)^{-\gamma}$$

a)  $N(\Delta t = 1 \text{ day}, \Omega = 1.7 \text{ sr}, A = 0.5625 \text{ m}^2, \epsilon \geq 10^{14} \text{ eV}) = \int_A \int_{\Omega} \int_{E=10^{14} \text{ eV}}^{\infty} \phi_1 \left( \frac{\epsilon}{E_1} \right)^{-\gamma} d\epsilon d\Omega dA$

$$N = \Omega A \Delta t \phi_1 \left[ \left( \frac{\infty}{E_1} \right)^{-\gamma+1} - \left( \frac{E_0}{E_1} \right)^{-\gamma+1} \right] \times \frac{E_1}{-\gamma+1}$$

$\gamma = 2.7 \Rightarrow \boxed{N = \Omega A \Delta t \phi_1 \frac{E_1}{\gamma-1} \left( \frac{E_0}{E_1} \right)^{1-\gamma}}$

$$\Delta t = 1 \text{ day} = 86400 \text{ s}$$

$$\phi_1 = 1.1 \times 10^{-18} \text{ eV} / \text{m}^2 / \text{sr} / \text{s}$$

$$E_1 = 10^{14} \text{ eV}$$

$$\boxed{N = 9.7}$$

b)

$$E_0 = 10 \text{ GeV} = 10^{10} \text{ eV}$$

$$A_{\oplus} = 4\pi R_{\oplus}^2 = 5.11202 \times 10^{14} \text{ m}^2$$

$$\Delta t = 1 \text{ s}$$

$$\Rightarrow \frac{dN}{dt}(\epsilon \geq 10 \text{ GeV}) = 2.6 \times 10^{18} / \text{s}$$

$m_p$

$$\Rightarrow t = 1 \text{ yr} \Rightarrow 3.8 \times 10^{-4} \text{ g}$$

$$t = 4.54 \text{ Gyr} \Rightarrow 1.7 \times 10^6 \text{ g} \sim 1.7 \text{ Ton.}$$

[c] convert to GeV !!

$$\phi(E) = 1.1 \cdot 10^{-18} \left( \frac{E}{10^{10} \text{ eV}} \right) / \text{GeV m}^2 \text{ s sr}$$

↓

$$\phi(E_{\text{in GeV}}) = 1.1 \cdot 10^{-9} \left( \frac{E}{10^9 \text{ GeV}} \right) / \text{GeV m}^2 \text{ s sr}$$

[d]

$$E^2 \phi(E) = E^2 \frac{dN}{dE dA d\Omega dt}$$

$$N(E) = \int_{E_0}^{\infty} \phi(E) dE = \frac{A \Delta t}{\gamma - 1} \phi_0 \left( \frac{E_0}{E_1} \right)^{1-\gamma}$$

$$\Rightarrow \frac{I(E)}{A \Delta E} = \frac{4\pi}{\gamma - 1} \frac{E_1}{E_0} \left( \frac{E}{E_1} \right)^{1-\gamma}$$

$$\frac{E_1}{\gamma - 1} \phi_0 \left( \frac{E_1}{E_0} \right)^{-\gamma}$$

$$N(E) = \frac{A \Delta t}{(\gamma - 1) E_1} \frac{E_1^2 \phi_0 \left( \frac{E_1}{E_0} \right)^{-\gamma}}{E^2 \phi(E)}$$

$$I(E) = \frac{A \Delta t}{\gamma - 1} \int_{E_0}^{\infty} \frac{E^{2-\gamma}}{E_0^{\gamma-1}} dE$$

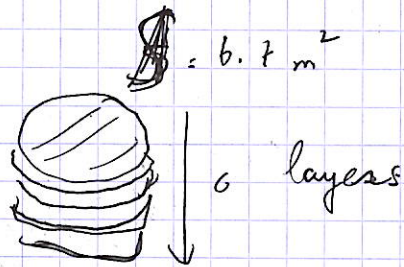
$$= \frac{A \Delta t}{\gamma - 1} \int_{E_0}^{\infty} E^{2-\gamma} dE$$

$$\Rightarrow \left[ \frac{E^{3-\gamma}}{3-\gamma} \right]_{E_0}^{\infty}$$

$$\Rightarrow \left[ \frac{E^{0.3}}{0.3} \right]_{0.3}^{\infty}$$

$$\Rightarrow \boxed{E^2 \phi(E) = \frac{N(\gamma - 1) E}{A \Delta t}}$$

## 2. Direct detection of CRs



design (1)

superconducting magnet

length 1m

$$B_1 = 0.75 \text{ T}$$

$$\Delta t_1 = 3 \text{ yrs}$$

design (2)

Permanent magnet

length L

$$B_2 = 0.15 \text{ T}$$

$$\Delta t_2 = 15 \text{ yrs.}$$

(a)

$$\frac{\sigma(p)}{P} = \frac{\sigma_{\text{pos}}}{\sqrt{N} L^2 B} P$$

$$\left( \frac{\sigma(p)}{P} \right)_1 = \left( \frac{\sigma(p)}{P} \right)_2 \Rightarrow \frac{1}{L_1^2 B_1} = \frac{1}{L_2^2 B_2}$$

$$\Rightarrow L = L_1 \sqrt{\frac{B_1}{B_2}}$$

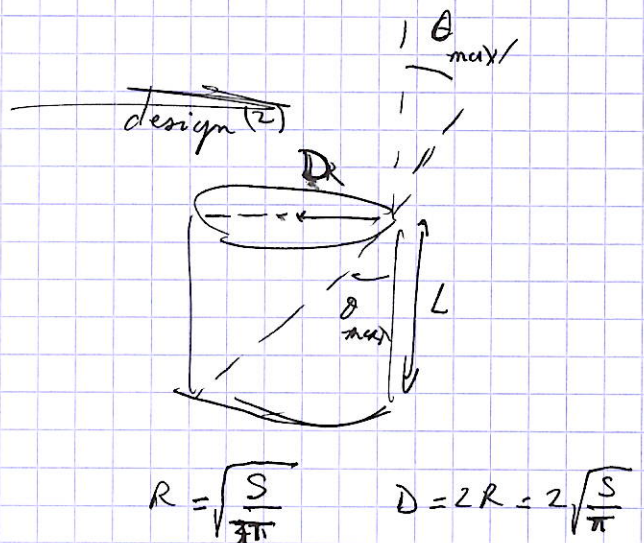
$$L = 1 \sqrt{\frac{0.75}{0.15}} = 2.24 \text{ m}$$

(b) ~~design (1)~~

$$A = S \pi \sin^2 \theta_{\text{max}}$$

$$\sin \theta_{\text{max}} = \frac{D}{\sqrt{D^2 + L^2}}$$

$$= \frac{D}{\sqrt{\frac{4S}{\pi} + L^2}}$$



$$\text{aperture} : A = \frac{S \pi}{\sqrt{\frac{4S}{\pi} + L^2}} \approx \sqrt{\frac{S}{\pi}}$$

design (1)

$$A = 18.84 \text{ m}^2 \text{ sr}$$

design (2)

$$A = 13.3 \text{ m}^2 \text{ sr}$$

from Sullivan.

$$G = \frac{1}{2} \pi^2 \left[ R_1^2 + R_2^2 + L^2 - \sqrt{(R_1^2 + R_2^2 + L^2)^2 - 4R_1^2 R_2^2} \right]$$

In our case  $R_1 = R_2 = R$

$$\Rightarrow G = \frac{1}{2} \pi^2 \left[ 2R^2 + L^2 - \sqrt{(2R^2 + L^2)^2 - 4R^4} \right]$$

$$R = \sqrt{\frac{S}{\pi}}$$

design 1:

$$G = \frac{31.86}{10.75} \text{ m}^2 \text{ sr}$$

design 2:

$$G = \frac{11.19}{5.13} \text{ m}^2 \text{ sr}$$

$$\text{Exposure: } \mathcal{E} = G \Delta t$$

$$\Delta t = 3 \text{ yrs}$$

$$\Delta t = 15 \text{ yrs.}$$

$$\mathcal{E}_1 = \frac{31.86}{10.75} 10^9 \text{ m}^2 \text{ sr s}$$

$$\mathcal{E}_2 = \frac{11.19}{5.13} 10^9 \text{ sr s m}^2$$