## Astroteilchenphysik II – Cosmic Rays

Please submit via ILIAS by Monday January 30th. Solutions will be discussed Feb. 2nd 9:45 in room 10/1.

## 1. Hadronic Particle Production

In general, 'energy scaling' of particle production means that the probability of producing particle A in reactions of particle B depends only on the ratio  $x = E_A/E_B$ ,  $E d\sigma/dE_A = f(x)$  (examples: A=photons from the bremsstrahlung of B=electrons or A=neutral pions from the collision of B=protons with protons). Please show that if the energy distribution of the projectiles B follows a power law  $E^{-\alpha}$ , then the energy distribution of produced particles A is a power law as well with same spectral index  $\alpha$ . (3 points)

## 2. Charged Pion Decay

Consider the decay  $X^{\pm} \rightarrow \mu^{\pm} + \overleftarrow{\nu}_{\mu}$ , where *X* is a meson of mass  $M_X$  moving with speed  $\beta_X$ .

- (a) Derive the energy distributions of neutrinos and muons,  $dn/dE_{\nu}$  and  $dn/dE_{\mu}$ , in the laboratory frame. Assume an isotropic distribution of particle emission in the rest frame of the meson.(2 points)
- (b) What is the average fraction of energy with respect to the meson's laboratory energy,  $z_{\nu/\mu} = E_{\nu/\mu}/E_X$ , carried away by neutrinos and muons in the limit  $\beta_X = 1$ ? Please give numerical values for  $z_{\nu/\mu}$  for charged  $\pi$ , K and D mesons. (1 point)

## 3. Atmospheric depth

Consider a primary cosmic-ray particle which, upon entering the Earth's atmosphere, interacts with a nucleus of air and produces secondary particles. For the primary proton and primary iron nucleus, crossing the atmosphere with a zenith angle of 45°, solve the assignments given below. The interaction lengths of protons and iron nuclei in air are  $\lambda_p = 80 \text{ g/cm}^2$  and  $\lambda_{Fe} = 12 \text{ g/cm}^2$ .

- (a) On average, at which atmospheric depth (measured in  $g/cm^2$ ) is the first interaction taking place? (1 point)
- (b) A detector is flown with a balloon at large heights so that the remainder of the atmosphere above the detector amounts only to 5.5 g/cm<sup>2</sup> of vertical column density. Calculate the fraction of protons and iron nuclei which interact at least once before reaching the detector. (1 point)
- (c) What is the mean height of the first interaction if you assume an isothermal atmosphere with the scale height of 8.4 km? (2 points)